

cated ways) and presents the elements of the subject in a simple, sensible form. Now that he has done it, we can only wonder why the book was so long in coming.

The book begins by examining the simple model of a finite-state machine (a machine which can have only a finite number of states). After carefully defining his mathematical model and explaining some of the relationships between it and reality, Minsky proves that a finite-state machine can add arbitrarily long sequences of binary digits of two binary numbers provided the lowest-order digits are presented to the machine first, but that a finite-state machine *cannot* do the corresponding multiplication.

Neural nets and the McCulloch-Pitts models are shown to be equivalent to automata made up of simple parts and then to be equivalent to finite-state machines. Some logical systems such as Kleene's theory of regular sets are also shown to be equivalent to finite-state machines; thus the idea of finite-state machines is shown to provide a unifying view of a number of originally diverse topics.

The second part of the book treats infinite-state machines. It begins with the development of the classic Turing machine and effective computability, demonstrates the usual universal Turing machine, and passes on to the surprising limitations of effective computability, such as the unsolvability of the "halting problem" and some decision problems. Special attention is given to the interesting consequences of computable numbers.

Just as finite-state machines were shown to unify some diverse topics, so infinite-state machines are shown to be related to recursive-function theory and Post's theory of productions, and the normal-form theorem is proven. The book ends with some specially simple Turing machine models such as the four-symbol, seven-state Turing machine.

These topics, many of which were until quite recently considered to be advanced and abstruse, are treated with a commonsense attitude, a simplicity of presentation, and a style that make much of this material accessible to the "average person" interested in the field. Thus the book is a significant contribution.

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Cool Stars

Colloquium on Late-Type Stars. Trieste, June 1966. MARGHERITA HACK, Ed. Osservatorio Astronomico, Trieste, 1967. 465 pp., illus. Paper, 10 L.

The term "late-type" is astronomical jargon for stars of low temperature. The idiom arises from a long-dead theory of stellar evolution that held that stars began their lives as objects with very high surface temperatures and then simply cooled down. To this day we call hot stars "early-type" and cool stars "late-type." The spectral sequence is similarly ordered, and the designation "early" or "late" refers to the position of the star in the sequence. More particularly, "late-type stars" refers to stars cooler than the sun, with surface temperatures ranging downward from about 5000°K, although no precise definition has ever been adopted. A number of objects are cooler than the reddest stars visible to the eye. Such objects are detected by infrared devices and are now often called "dark-brown" stars.

Cool stars are of astrophysical interest for a variety of reasons: At low temperatures, compounds form. Isotope effects in molecular bands of carbon compounds such as CN are readily observed, for example. Only from molecular spectra has it been possible to get information on stellar isotope abundances. The spectra of cool stars are very sensitive to the ratio of carbon to oxygen. Titanium oxide bands dominate the spectra of cool stars with solar composition, whereas if carbon exceeds oxygen in abundance all the oxygen disappears as carbon monoxide and carbon bands dominate the spectra. Giants and supergiants represent stars in late stages of evolution; cool representatives sometimes show remarkable abundance anomalies and even the presence of short-lived technetium. Hence these objects are of enormous interest in connection with theories of element building and stellar structure. A great advantage in their study is offered by the fact that the astrophysicist can approximate their temperatures in the laboratory.

This excellent symposium volume brings together a number of valuable papers on the spectral classification, high-dispersion spectra, chemical and isotopic composition, atmospheres, internal structures, and evolution of these remarkable stars. As mentioned above, some of them are very luminous giants

and supergiants; others are dwarf stars much fainter than the sun and show no evolutionary effects, but attempts have been made to find planet-like companions among them. Perhaps the greatest value of this information-packed book is the stimulation it will certainly bring to a rising generation of physicists, chemists, and astrophysicists to engage in one of the most exciting fields in astronomy.

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The Views of a Physicist

Erwin Schrödinger. An Introduction to His Writings. WILLIAM T. SCOTT. University of Massachusetts Press, Amherst, 1967. xvi + 175 pp. \$6.50.

In his later years Erwin Schrödinger was deeply concerned that physics was cutting itself off from its historical background. He considered much in quantum mechanics, particularly the statistical interpretation of the wave function, to be a break with the scientific tradition of over two millennia. Schrödinger argued as forcefully as he could against this trend, going so far as to compare quantum jumps and virtual transitions with the circles and epicycles of medieval astronomy. But this was the same Schrödinger who had created wave mechanics in 1926 in a series of articles about which Max Born later remarked, "What is there more sublime in theoretical physics?" Schrödinger had indeed felt that same way then, being confident that his new theory "struck deep into the nature of the quantization rules" because it made atomic quantum numbers as *natural* as the number of nodes of a vibrating string. He had never liked "this damned quantum jumping," as he called it, and had found the essentially discrete and algebraic matrix mechanics to be "forbidding, not to say repulsive." He did not find in it what he prized most—an intuitive clarity.

William T. Scott has written the first book on this remarkable physicist. It is, in his own words, "a modest introduction" which deals only with some of the major themes in Schrödinger's writings and very briefly with his life. (His publisher does him a disservice by misadvertising the book as a "comprehensive study.") One of these major themes, strange in view of Schrödinger's attitude toward quantum mechanics, is the