

Reports

Xenon Fluorosilicate and Related Compounds

Abstract. *The reaction product of Xe, F₂, and SiF₄ in a glow discharge has been shown to have a composition approximating Xe₃SiF₆. It and a similar hexafluorophosphate are unstable at room temperature. A stable hexafluoroantimonate is formed from Xe, F₂, and SbF₅ at 250°C.*

The initial announcement of the preparation of a salt of xenon(I), XePtF₆ (1), was followed by reports of the preparation of XeRhF₆ by Bartlett (2) and of the "adduct" Xe-RuF₆ by Chernick *et al.* (3). These were prepared by combination of the noble gas with the noble metal hexafluoride. On the basis of the magnetic susceptibility the first of these was characterized as Xe^IPt^VF₆, although in the absence of electron-spin resonance data there appeared to be no good reason for preferring this formulation over Xe^IPt^{IV}F₆. More recently Bartlett (4) has provided more evidence for the Xe^IPt^VF₆ formulation by metathetical reactions in IF₅. Bartlett (2) has also reported the formation of brick red Xe(PtF₆)₂ by the action of Xe on excess solid PtF₆ or by pyrolysis of XePtF₆. More recently Peacock and co-workers (5) have reported the formation of yellow, diamagnetic XeF₂·2SbF₅ by reaction of XeF₂ or XeF₄ with SbF₅ and of straw-colored XeF₂·2TaF₅ from XeF₄ and TaF₅. These are suggested to be covalent compounds of xenon(II).

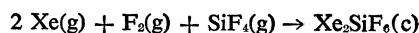
We have prepared a salt of xenon which appears with little doubt to be a compound of xenon(I). When a 2:1 mole ratio of xenon and fluorine at a total pressure of less than 10 mm was passed in glass apparatus at -78°C through a 6500- to 7000-volt glow discharge, a white crystalline solid was deposited on the walls of the vessel. This solid was stable at -78°C, but upon warming to room temperature slowly decomposed to gaseous products among which Xe and SiF₄ were found by mass spectral analysis. The gaseous decomposition products were

hydrolyzed in 10-percent aqueous KOH, and the silicon and fluorine content of the solution was determined by standard analytical methods. These, combined with the volume of xenon liberated, in a typical case gave a composition for the solid of Xe₃SiF₆. A similar run carried out with a 2:1:1 mole ratio of xenon, fluorine and silicon tetrafluoride gave a product having a 2.2:1 atom ratio of xenon to silicon. In the case of the 2:2:1 ratio about 0.35 g of product was obtained. Although it is possible to obtain molecular addition compounds under conditions such as these—as has, indeed, been suggested for the Xe-RuF₆ "adduct"—it is a little more difficult to imagine a ternary Xe-F₂-SiF₄ adduct of proportions so fortuitously close to the composition Xe₃SiF₆ and the formulation (Xe⁺)₃SiF₆⁻ seems to be the most reasonable.

A similar reaction with xenon, phosphorus pentafluoride, and fluorine produced a white-to-yellow solid, unstable at room temperature and containing both Xe and PF₅ as judged by mass spectral and infrared analyses of the decomposition products. On the other hand, heating a mixture of the gases at 220°C for 10 hours in a nickel bomb and passing the gases into a trap at -78°C produced nothing containing xenon.

However, when 0.153 mole of SbF₅, 0.0435 mole of Xe, and 0.0410 mole of F₂ were heated in a nickel bomb at 250°C for 12 hours, a xenon-containing, pale yellow solid resulted which reacted vigorously with water with release of xenon. This did not sublime below 100°C, whereas the compound reported by Peacock (5) sublimed at 60°C in vacuum. Our compound is presumed to be XeSbF₆.

If the radius of the Xe⁺ cation is similar to that for Na⁺ (6) and the lattice energy for Xe₃SiF₆ is similar to that for Na₃SiF₆ (7), the molar enthalpy for the reaction



is about -40 kcal. If it is assumed that the entropy of Xe₃SiF₆ is the same as that of Cs₃SiF₆, ΔS for the reaction

is about -134 entropy units and ΔF is about 0 at 298°K. Since the reverse of this reaction is the expected decomposition reaction, the calculations correspond to the qualitatively observed stability. Similar calculations indicate that XeBF₄ should be stable up to about 150°C, but that KrBF₄ should be unstable down to within a few degrees of absolute zero.

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References and Notes

1. N. Bartlett, *Proc. Chem. Soc.* **1962**, 218 (1962).
2. See p. 38, *Chem. Eng. News* **63**, 36 (1963).
3. C. Chernick *et al.*, *Science* **138**, 136 (1962); H. H. Claassen, J. G. Malm, H. H. Selig, Conference on Noble Gas Compounds (Argonne National Laboratory, 22-23 Apr. 1963).
4. N. Bartlett and N. K. Jha, *Noble Gas Compounds*, H. H. Hyman, Ed. (Univ. of Chicago Press, Chicago, 1963), p. 23.
5. A. J. Edwards, J. H. Holloway, R. D. Peacock, *Proc. Chem. Soc.* **1963**, 275 (1963).
6. The radii of the noble gas cations (for example, Xe⁺) were estimated from the ratio of the halogen van der Waals radii to the ionic radii calculated by Bell and Gelles, *J. Chem. Soc.* **1951**, 2734 (1951), and the van der Waals radii of the noble gases—2.07 Å for Xe, as determined from the crystal structure of XeF₂ [P. A. Agron, G. M. Begun, H. A. Levy, A. A. Mason, C. G. Jones, D. F. Smith, *Science* **139**, 842 (1963)].
7. There should be no crystal field stabilization energy for the noble gas + 1 cations as contrasted to the halogen + 1 cations [J. Arotsky and M. C. R. Symons, *Quarterly Reviews* **16**, 282 (1962)].

24 December 1963

Origin of Diamonds in the Ureilites

Abstract. *Diamonds in two of the three diamantiferous stony meteorites (ureilites) show a pronounced crystallographic orientation. This suggests that the diamonds were produced by shock during breakup of the meteorite's parent body rather than by high gravitational pressure.*

In previous papers (1, 2) on the origin of meteoritic diamonds (Table 1) I concluded that the diamonds in the Canyon Diablo iron were formed by conversion of graphite as a result of the shock of impact of the meteoroid with the earth. The subsequent finding of coesite and stishovite (3) in the sandstone around the Canyon Diablo crater indicates shock pressures of at least 100 kbar. The successful shock-induced conversion of graphite to diamond in the laboratory (4) also lends strong support to the shock origin of Canyon Diablo diamonds.

Diamonds also occur in a rare type