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eastern basin is 1.9 to 3.8×10^{12} g N year⁻¹. Finally, comparing our conservative estimate of dissolved inorganic nitrogen discharge from a flooding Nile $(0.06 \times 10^{12}$ g N year⁻¹) to total eastern Mediterranean new nitrogen production today (1.9 to 3.8×10^{12} g N year⁻¹) demonstrates that nutrient delivery from the Nile was unlikely to have increased new nitrogen production by more than 1.6 to 3.2% over contemporary rates. Increasing the Nile discharge rate beyond 2.5 times the modern value, and adding a contribution from Eurasian rivers such that riverine runoff was four times that of the predam Nile discharge would only increase new nitrogen production by 5 to 10%.

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Heavy Nitrogen in Carbonatites of the Kola Peninsula: A Possible Signature of the Deep Mantle

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Nitrogen and argon isotopes were measured in carbonatites and associated rocks from the Kola Peninsula in Russia. The Kola mantle source, which is thought to be located in the deep mantle, is enriched in heavy nitrogen (+3 per mil relative to air) as compared to Earth's surface (atmosphere and crust, +2 per mil) and the shallow mantle (-4 per mil). Recycling of oceanic crust (+6 per mil) or metal-silicate partitioning may account for the nitrogen isotopic composition of the deep mantle.

Determining the structure and composition of the mantle is necessary in order to understand current and past mantle dynamics and mantle-to-crust interactions. For example, some have argued that part of the mantle has been isolated from mantle convection for most of Earth's history, allowing the preservation of a primitive component that is only sampled by plumes. Evidence in support of this view arises mainly from noble gases, because plume-derived magmas often show lower radiogenic/primordial isotope ratios than do mid-ocean ridge basalts (MORBs) (1). At variance with these models are geophysical (2), experimental (3), and geochemical (4) lines of evidence that suggest that some of the subducting slabs sink through the 670-kmdeep seismic discontinuity, implying global stirring of the mantle.

The nitrogen isotopic composition of the shallow mantle that feeds MORs, which is expressed as per mil (‰) deviation relative to the composition of air [in δ^{15} N notation (5)], is estimated to be ~-4‰ (6-13). This signature

may be a remnant of the nitrogen isotopic composition of Earth-forming planetesimals, which later evolved as a result of (i) addition of meteoritic (14, 15) and cometary (15) volatiles, or (ii) fractional loss of atmospheric volatiles (16), or both. Biologic activity leads to a fractionation of nitrogen isotopes and gives sedimentary rocks a specific composition $[\delta^{15}N \sim +6\%]$ (17)]. Nitrate is used by denitrifying bacteria as the terminal electron acceptor in energy generation when oxygen is unavailable. Associated with denitrification is a kinetic isotope effect that enriches the residual nitrate in ^{15}N (18). Because NO₃⁻ is the main nitrogen-bearing nutrient, marine organisms and sediments are enriched in ^{15}N relative to the atmosphere (17). The oceanic crust thus enriched in ¹⁵N is subducted back into Earth at convergent plate margins, which makes nitrogen a potentially powerful tracer of volatile recycling in the mantle.

Mantle plumes, which are assumed to be fixed relative to plate motion, sample a deeper region than that feeding MORs (19). The isotopic composition of nitrogen in plumes has not been documented (20). Low ⁴He/³He ratios and a steep 20 Ne/²²Ne-²¹Ne/²²Ne correlation (21) indicate that a mantle plume contributed to the 370-million-year-old (22) ultrabasic-alkalinecarbonatite magmatism in the Kola Peninsula in Russia (the eastern segment of the Baltic Shield). The Kola rocks crystallized at depth (23), which prevented them from being extensively outgassed, minimizing any subsequent

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contamination by atmospheric volatiles. To document the nitrogen isotopic composition of the deep, plume-type mantle, we have measured nitrogen and argon isotopes in carbonatites and associated rocks from the Kola Peninsula [referred to as Kola rocks hereafter (Table 1)]. Volatiles trapped in fluid inclusions were released by stepwise crushing, followed by online nitrogen-argon purification and high-resolution static-vacuum mass spectrometry (10, 24).

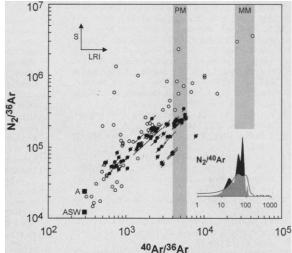
The Ne isotopic composition of the mantle is different from that of the atmosphere (25), making this element a Rosetta stone when deciphering the isotopic message of terrestrial volatiles. Neon and Ar isotopic ratios correlate in Kola rocks and MORBs, which allows us to estimate the deep and shallow mantle 40 Ar/ 36 Ar ratios at 4000 to 6000 (21) and at 25,000 to 44,000 (26), respectively (40 Ar is a decay product of 40 Kr.

Fig. 1. N₂/³⁶Ar-⁴⁰Ar/³⁶Ar correlations in MORBs [white dots (10, 13, 27)] and Kola rocks [black dots (Table 1)]. Data points fall on mixing lines between the mantle (PM, plumetype deep mantle; MM, MORtype shallow mantle) and Earth's surface (A, air; ASW, air-saturated water; S, sediments). This correlation was lately disturbed by ⁴⁰Ar radiogenic ingrowth (LRI). The uncertainties (depicted as the main axis of standard ellipses) are correlated because they mainly result from the ³⁶Ar blank correction. For the purpose of clarity, uncertainties for MORBs are omitted. The source end-member ⁴⁰Ar/³⁶Ar ratios (shaded regions) are derived from Ar-Ne isotope correlations

whereas ³⁶Ar is primordial in origin). These ratios are used hereafter to determine the isotopic and elemental characteristics of the mantle.

In Kola rocks [and MORBs (10, 13, 27)], the $N_2/{}^{36}Ar$ ratio correlates with the ${}^{40}Ar/{}^{36}Ar$ ratio (Fig. 1). The 40 Ar/ 36 Ar ratio is a reliable indicator of shallow contamination by sediments, air, or air-saturated water. A comparatively K-rich $(3.59 \times 10^{-4} \text{ mol g}^{-1})$ ijolite exhibits systematically high ⁴⁰Ar/³⁶Ar ratios at low $N_2/^{36}Ar$ ratios, which may reflect a late (postcrystallization) ⁴⁰Ar radiogenic ingrowth. A sedimentary contribution is suggested by the high $N_2/^{40}$ Ar ratios (up to 114) of some carbonatites. During partial melting, vesiculation, and formation of fluid inclusions, one might expect a complete extraction of N and Ar and an inheritance of the isotopic and elemental compositions of the source (28). When extrapolating the mixing relationships to the source end-member ${}^{40}\text{Ar}{}^{36}\text{Ar}$ ratios, one can estimate the N₂/ ${}^{36}\text{Ar}$ ratios of the deep mantle (from the Kola rocks) and shallow mantle (from the MORBs) at approximately 3×10^5 and $5 \times$ 10^6 , respectively. The N₂/ ${}^{40}\text{Ar}$ ratio of the deep mantle (60 ± 5) is lower than that of the shallow mantle [124 ± 20 (13)], which leads to a new estimate of the bulk N₂ content of the mantle ($3.6 \pm 1.4 \times 10^{-8} \text{ mol g}^{-1}$), as derived from a global K-Ar-N mass balance of the silicate Earth (27).

As in the case of MORBs (10, 13), Kola rocks form a triangular array in a N-Ar isotope plot (Fig. 2). This feature is consistent with mixing between air (or air-saturated water), sediments, and mantle gases. The fluid inclusions that host the heavy nitrogen show plume-type 4 He 3 He and 21 Ne ${}^{/2}$ Ne ratios and comparative-



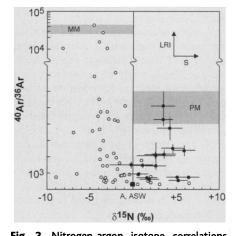


 Fig. 2. Nitrogen-argon isotope correlations. Note the change in scale at ⁴⁰Ar/³⁶Ar = 10⁴. Symbols and abbreviations are as in Fig. 1. MORBs (10, 13) and Kola rocks (Table 1) form triangular mixing arrays between the mantle, air (or air-saturated water), and sediments.

(21, 26). (Bottom right) Kernel density estimate (36, 37) of the $N_2/^{40}Ar$ ratios in MORBs (black area) and Kola rocks (white area).

Complex	Rock (mineral)	N ₂ (10 ⁻⁹ mol g ⁻¹)	δ ¹⁵ N (‰)	⁴⁰ Ar (10 ⁻¹¹ mol g ⁻¹)	⁴⁰ Ar/ ³⁶ Ar	K (10 ⁻⁵ mol g ⁻¹)
Lesnaya Varaka	Dunite with perovskite and titanomagnetite	3.96 ± 0.10	1.9 ± 1.1	5.68 ± 0.14	612 ± 55	<0.2
	Dunite (magnetite)	1.74 ± 0.05	3.5 ± 1.4	3.89 ± 0.13	5103 ± 899	<0.2
Kovdor	Turjaite	6.06 ± 0.15	3.5 ± 1.1	13.02 ± 0.33	4281 ± 394	3.2 ± 0.3
	Dunite	1.42 ± 0.04	4.6 ± 1.3	2.63 ± 0.07	2497 ± 266	5.1 ± 0.5
	Phoscorite (magnetite)	1.30 ± 0.03	6.0 ± 1.3	2.57 ± 0.06	2376 ± 308	<0.2
	Pyroxenite (diopside)	0.27 ± 0.01	2.7 ± 2.2	0.42 ± 0.01	2062 ± 820	6.2 ± 0.6
Seblyavr	Pyroxenite	3.40 ± 0.09	2.6 ± 1.2	3.93 ± 0.10	2102 ± 209	15.3 ± 0.3
	Pyroxenite (pyroxene)	0.57 ± 0.01	-0.2 ± 2.5	0.63 ± 0.02	1485 ± 201	12.1 ± 0.2
	Calcite carbonatite	7.96 ± 0.20	5.3 ± 1.1	7.00 ± 0.19	532 ± 51	17.6 ± 0.4
	Calcite carbonatite	20.20 ± 0.51	1.7 ± 1.2	27.23 ± 0.69	691 ± 61	0.42 ± 0.04
	Dunite	6.08 ± 0.15	3.6 ± 1.4	5.27 ± 0.14	2144 ± 251	13.4 ± 0.3
	Dolomite carbonatite	2.50 ± 0.06	6.5 ± 1.2	3.38 ± 0.09	745 ± 89	1.5 ± 0.1
	Ijolite	1.75 ± 0.04	4.3 ± 1.5	4.68 ± 0.12	3720 ± 667	35.9 ± 0.7
	Apatite-phlogopite-diopside ore	1.49 ± 0.04	2.6 ± 1.2	2.10 ± 0.05	1415 ± 166	10.2 ± 1.0
	Apatite-diopside-phlogopite ore	6.91 ± 0.17	1.1 ± 1.2	9.23 ± 0.23	1456 ± 132	26.3 ± 0.5
	Clinopyroxenite with perovskite and titanomagnetite	3.52 ± 0.09	0.7 ± 1.1	5.10 ± 0.13	926 ± 86	2.8 ± 0.3
	Dolomite carbonatite	2.83 ± 0.07	5.1 ± 1.2	3.33 ± 0.08	716 ± 68	0.21 ± 0.02

Table 1. N-Ar-K data in carbonatites and associated paragenesis from the Kola Peninsula (additional data are available at www.sciencemag.org/ feature/data/1044003.shl). Uncertainties are one sigma.

ly high 40Ar/36Ar ratios, which all together indicate a mantle source region (21). The coupled Sr and Nd isotope systematics in the Kola Peninsula provide no evidence for crustal contamination (29). Independently, preliminary measurements in our laboratory on basalt glasses from the Society Island hot spot yield positive δ^{15} N values, suggesting that heavy nitrogen is a common feature of plume-derived magmas. In contrast to the characteristics of the shallow mantle, which is depleted in heavy nitrogen by $\sim -4\%$ relative to the atmosphere (6–13), the deep mantle appears to be enriched by $\sim +3\%$. The δ^{15} N value of the deep mantle outside the range defined by the shallow mantle (6-13) and Earth's surface [crust and atmosphere, $\sim +2\%$ (10)] suggests that the heavy nitrogen is not primordial in origin. It is noteworthy that the late accretion of extraterrestrial matter depleted in ¹⁵N relative to the deep mantle would fail to account for the 15N depletion of the shallow mantle relative to the atmosphere.

Long-term isotopic fractionation of N during magma genesis and subsequent degassing, if any (28), are unable to account for the isotopic heterogeneity of present-day Earth. Models of N isotope fractionation between silicate melt and vapor predict that the residual N in the silicates after magma degassing should be enriched in ¹⁵N (7, 30). Because noble gas isotopic ratios suggest that the deep mantle is less degassed than the shallow mantle (1), one should expect the shallow mantle to be enriched in ¹⁵N relative to the deep mantle, which is contrary to the observation.

Nitrogen is a siderophile element (31), whereas Ar is not (32). Hence, uptake of N by the core would lower the N_2 /³⁶Ar ratio of the mantle. The solubility of N_2 in Fe-Ni alloys follows Sieverts' law (N_2 gas \rightleftharpoons 2N dissolved). Because the ¹⁴N-¹⁴N bond is weaker than the ¹⁵N-¹⁴N or ¹⁵N-¹⁵N bonds, the residual N in the mantle is presumably enriched in ¹⁵N relative to its metal counterpart. Such a possibility would require appreciable fractionation of nitrogen isotopes to account for the isotopic stratification of present-day Earth, which calls for experimental as well as theoretical confirmation.

Phanerozoic marine sediments and the deep mantle both exhibit high $N_2/^{36}$ Ar ratios and high $\delta^{15}N$ values relative to the atmosphere. An appealing possibility is that the nitrogen carried from the deep mantle to the surface of Earth by plumes is of recycled origin. This interpretation is consistent with the recycling of Ar in Earth's interior (21, 33) but creates a paradox. If the ¹⁵N depletion of the shallow mantle was primordial in origin, as is generally thought (6-16), preferential recycling of sediments in the deep mantle would at the same time increase the $N_2/$ ³⁶Ar ratio and the δ^{15} N value of the deep mantle relative to the shallow mantle, which is apparently not the case. A way to get around this difficulty is to consider that the nitrogen isotopic composition of the shallow mantle is not primordial in origin but rather results from the recycling of Archean, ¹⁵N-depleted, sedimentary organic matter (34). An alternative explanation is that the $N_2/^{36}$ Ar ratio of the primitive mantle was lower than the inferred present-day ratio of the deep mantle $(\sim 3 \times 10^5)$. Early degassing of the shallow mantle would have increased the $N_2/^{36}Ar$ ratio to $\sim 5 \times 10^6$ as a result of preferential retention of nitrogen in silicates under the reducing conditions prevailing at that time (35). Later on, the $N_2^{/36}$ Ar ratio of the deep mantle would have increased to its presentday value through recycling of surface nitrogen. Further investigation of the nitrogen isotopic composition of ocean islands such as Hawaii or Iceland will allow this unique stable isotope heterogeneity of Earth to be documented.

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- 24. Samples weighing 0.1 to 0.4 g were crushed under ultrahigh vacuum (~2 \times 10 $^{-9}$ torr) with a soft iron piston cylinder activated by an external solenoid. Gases were then purified in a glass line where H, C, and S compounds were oxidized (in the presence of Pt foils held at 723 K) with O2 produced by heating a CuO furnace at 1023 K (4CuO \rightleftharpoons 2Cu₂O + O₂). Condensible gases (for example, H2O, CO2, and SO2) were trapped in a cold finger held at 90 K. Excess oxygen was resorbed onto the CuO at 723 K. Nitrogen was then admitted into a Micromass VG5400 static-vacuum mass spectrometer. The trap current of the Nier type source was 20 μ A, resulting in a $\rm N_2$ sensitivity of 2.8 \times 10^{-5} A torr^{-1}. The nitrogen isotopic composition was determined as the 29/28 ion beam ratio on a Faraday collector. Possible isobaric interferences with N2 of CO, N2H, and C2H2 at masses 28 and 29 were monitored at mass 30 and were always negligibly small. Gases left in the purification line were further purified over hot Ti sponge getters. Argon was then admitted to the mass spectrometer, where ⁴⁰Ar was analyzed with a Faraday collector (with a sensitivity of 5.0 \times 10⁻⁵ A torr⁻¹) and ³⁶Ar was measured by ion counting after amplification by an electron multiplier.
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