

too small to explain a fluctuation of ~ 0.25 MAR. A large century-scale imbalance for the Antarctic interior is unlikely. This conclusion is in keeping with a body of relative sea-level and geodetic evidence supporting the notion that the grounded ice has been in balance at the millennial scale (31).

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25. Given data are 0.15 MAR "short-term variability" (20). Values calculated from tabulated data are 0.03 MAR (23) and 0.11 MAR (17); values calculated from graphic data are 0.14 MAR (19), 0.17 MAR (21), and 0.19 MAR (18).
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28. An ice imbalance of -0.28 MAR is $-3.9 \text{ cm year}^{-1}$ in the ROC. The elevation rate is $-0.9 \text{ cm year}^{-1}$, leaving 3.0 cm year^{-1} of snowfall fluctuation. This is equivalent to $3.0\rho_{\text{snow}}/\rho_{\text{ice}} = 1.15 \text{ cm year}^{-1}$ of ice, or 0.08 MAR (where ρ_{ice} is the density of ice).
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Migration of Fluids Beneath Yellowstone Caldera Inferred from Satellite Radar Interferometry

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Satellite interferometric synthetic aperture radar is uniquely suited to monitoring year-to-year deformation of the entire Yellowstone caldera (about 3000 square kilometers). Sequential interferograms indicate that subsidence within the caldera migrated from one resurgent dome to the other between August 1992 and August 1995. Between August 1995 and September 1996, the caldera region near the northeast dome began to inflate, and accompanying surface uplift migrated to the southwest dome between September 1996 and June 1997. These deformation data are consistent with hydrothermal or magmatic fluid migration into and out of two sill-like bodies that are about 8 kilometers directly beneath the caldera.

Yellowstone National Park (Fig. 1) is famous for its numerous hydrothermal features and other natural wonders but is better known among earth scientists as the site of the world's largest restless caldera. Many earth scientists believe the park is the present-day terminus of the active Yellowstone hotspot. The hotspot track can be traced along a string of large calderas (1) to its origin ~ 16 million years ago in southeastern Oregon and northern Nevada. The three most recent caldera-forming events occurred during cataclysmic rhyolite ash-flow tuff eruptions during the past 2.1 million years in the Yellowstone region (1, 2). Yellowstone caldera,

the youngest of the three (Fig. 1), formed $\sim 630,000$ years ago in an eruption (many times larger than any historic volcanic eruption) that ejected $\sim 1000 \text{ km}^3$ of debris—about 1000 times the volume of magma erupted at Mount St. Helens in May 1980 (1). A subsequent episode of dominantly extrusive volcanism buried Yellowstone caldera under rhyolite lava flows from 150,000 to 70,000 years ago (2). Even though the last magmatic eruption occurred $\sim 70,000$ years ago, geologic and geophysical evidence suggests that a crustal magma reservoir beneath Yellowstone is maintained in a partly molten state by episodic intrusions of basaltic magma (3). Because another caldera-forming eruption is almost inevitable, though not imminent, a continuous monitoring program is important. It is equally important, however, to assess the patterns of deformation

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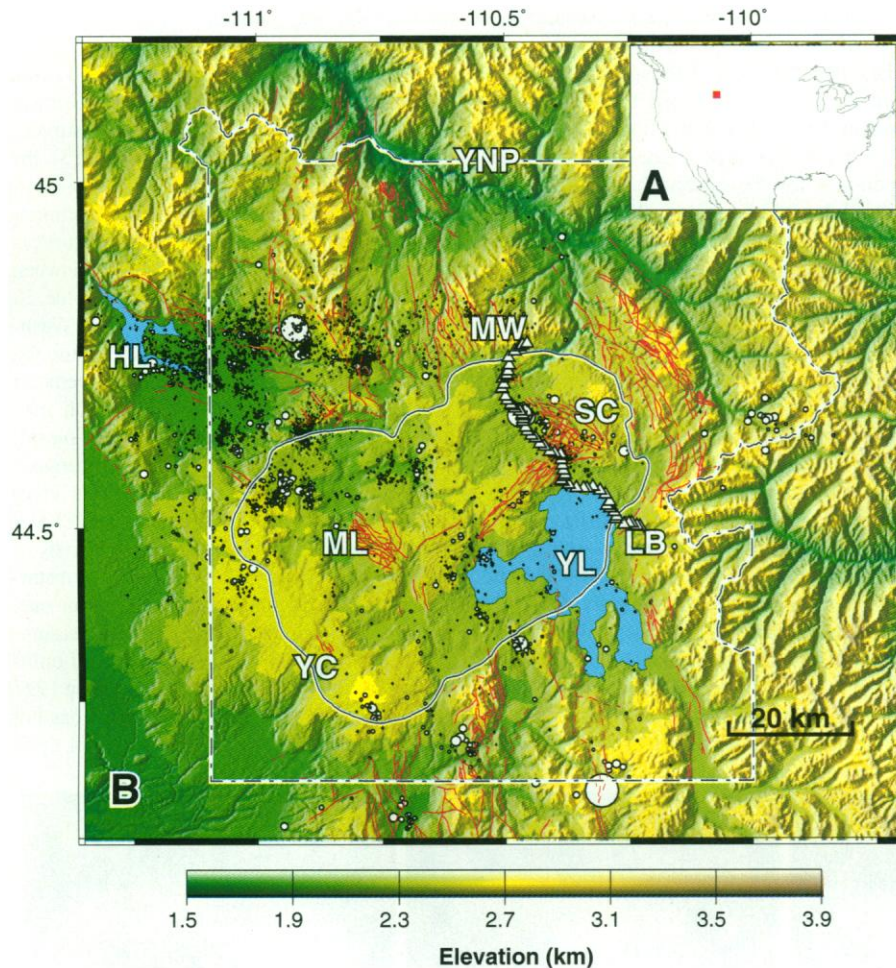


Fig. 1. Location maps for this study. (A) Regional scale map. The location of the detailed study area is indicated by the red square. (B) Topographic map of detailed study area. The topographic data is derived from USGS DEMs. For reference and discussion, the most important physical and geographic features are labeled as follows: HL, Hebgen Lake; LB, Lake Butte; ML, Mallard Lake dome; MW, Mount Washburn; SC, Sour Creek dome; YC, Yellowstone Caldera boundary (solid black line); YL, Yellowstone Lake; YNP, Yellowstone National Park boundary (dashed black line). The string of triangles running northwest from LB to MW marks the location of a leveling line to which displacement inferred from InSAR measurements is compared in Fig. 4. Red lines show faults mapped (solid) and inferred (dashed) (2). The epicenters of earthquakes occurring in the time interval between the acquisition dates of the earliest and latest radar images (Table 1) are shown as white-filled black circles. The size of a circle is proportional to the exponential of the magnitude of the earthquake. The smallest earthquakes shown are magnitude ~ 0.0 , whereas the three largest earthquakes shown are a magnitude 5.1 earthquake (depth 1.9 km, 28 August 1995) off the south edge of YNP, a magnitude 4.9 earthquake (depth 5.7 km, 26 March 1994) ~ 20 km east of HL, and a magnitude 4.8 earthquake (depth 3.2 km, 24 September 1994) on the west edge of SC. The 4884 earthquakes plotted as black dots are taken from the University of Utah Seismographic Stations' Yellowstone National Park Earthquake Catalogs for 1983–1992.

of the caldera in an effort to understand the magmatic plumbing beneath Yellowstone caldera and large calderas in general.

Field studies have revealed that Yellowstone has experienced uplift and subsidence of its caldera floor in historic and prehistoric times. Deformed postglacial terraces around Yellowstone Lake (YL) (4) and geomorphic evidence for changes in the Yellowstone River gradient (5) suggest multiple episodes of uplift and subsidence within the caldera in the past 12,000 years. Uplift of the caldera floor in the 20th century was documented by comparing leveling surveys made in 1923 and 1975–1977 (6). Re-leveling of benchmarks within the caldera along a single leveling line (Fig. 1) has been conducted on a near-yearly basis from 1983 through 1995 (7–10). Results show that the northeast caldera floor experienced uplift from 1923 through 1984, no measurable deformation in 1984–1985, and subsidence from 1985 through 1995. Less frequent trilateration and GPS (Global Positioning System) measurements (8, 11) are consistent with the leveling results and imply depressurization of a relatively shallow source beneath the caldera during the recent period of subsidence.

The main advantage of satellite interfero-

metric synthetic aperture radar (InSAR) is its ability, under favorable conditions, to measure deformation of the entire caldera floor and its surroundings. By measuring a surface of deformation rather than movement at isolated points, we are able to locate and characterize deformation sources more precisely than has been possible to date. High-resolution synthetic aperture radar (SAR) images of Yellowstone were formed with the use of raw satellite radar data from the European Space Agency. The interferograms (Fig. 2) were

made by taking the phase difference of pairs of SAR images, correcting for the topography [for example, see (12)] using a digital elevation model (DEM) (Fig. 1), and then correcting for orbital errors (13). We selected radar images acquired only during the summer months to minimize the decorrelation between images resulting from snow accumulation. The summer images were then combined (Table 1) to pick paired images that minimized the effects of uncorrected topography [resulting from inaccuracies in the

Table 1. Information on the individual satellite images used to construct the six interferograms shown in this report. The phase differences are relative to the master image. The altitude of ambiguity h_a (25), determined by the separation in orbits, is the amount of topographic change required to generate one interferometric fringe.

Fig. No.	Master image			Slave image			h_a (m)	Frame/track
	Orbit No.	Satellite	Acquisition date (year/month/day)	Orbit No.	Satellite	Acquisition date (year/month/day)		
2A	5697	ERS1	92/08/17	10206	ERS1	93/06/28	–213	2709/41
2B	10206	ERS1	93/06/28	21572	ERS1	95/08/30	–54	2709/41
2C	21572	ERS1	95/08/30	7410	ERS2	96/09/19	80	2709/41
2D	5697	ERS1	92/08/17	20570	ERS1	95/06/21	102	2709/41
2E	5196	ERS1	92/07/13	21572	ERS1	95/08/30	–85	2709/41
2F	20849	ERS1	95/07/11	11196	ERS2	97/06/11	78	887/320

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DEM (14)] and the possibility of baseline decorrelation [resulting from a large separation in the orbits (15)].

The interferograms (Fig. 2) indicate that deformation is episodic on varying time scales. On a time scale of ~ 1 to 2 years, the center of deformation changes within the caldera from the Sour Creek (SC) dome (Fig. 2, A and C) to the Mallard Lake (ML) dome (Fig. 2, B and F). On a longer time scale of ~ 3 years (Fig. 2, D and E, and Table 1), the entire caldera floor appears to subside. The initiation of renewed uplift (Fig. 2, C and F) represents the end of ~ 10 years of caldera subsidence. The progression of uplift from the SC dome to the ML dome may indicate that SC is closer, or at least better connected, to the magmatic source driving the deformation. The apparent lateral migrations of deformation shown in Fig. 2, A through C and F, are probably not artifacts of atmospheric disturbances in the radar images, because the disturbances would need to move from one resurgent dome to the other in both pairs of successive interferograms. Atmospheric disturbances take the form of spatially variable atmospheric water vapor content and possibly spatially variable electron content in the

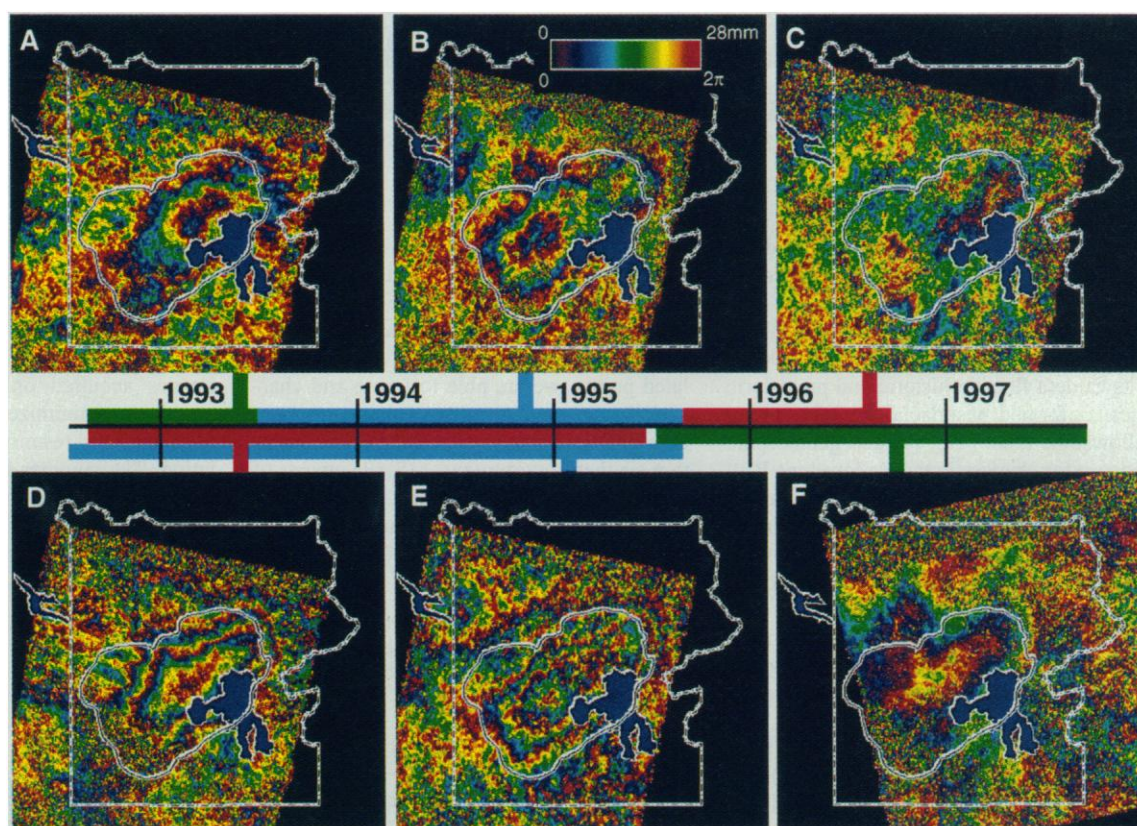
ionosphere. There is no reason to suspect that either of these effects should first favor one dome, then the other, although these effects could have locally distorted the deformation signal in the interferograms (Fig. 2).

To estimate the depth of the deforming body responsible for the observed subsidence, we modeled the interferogram in Fig. 2B with two tabular sill-like bodies (16). We chose to model Fig. 2B because it has the best combination of a long time interval and good overall image coherence. We also chose to use two sill-like bodies because the pattern of fringes in Fig. 2B requires more than one source. The depth of the larger body in the model (Fig. 3A, better constrained than the small body) is 8.5 ± 4 km (95% confidence intervals). The depth coincides with that at which Miller and Smith (17) find low seismic velocity bodies beneath both resurgent domes. Previous studies placed the body at a depth of 10 ± 5 km (8) and 3 to 6 km [possibly extending to 9 km beneath ML (18)]. The model requires a decrease in volume of 0.035 to 0.057 km^3 from June 1993 to August 1995 or a volumetric decrease rate of 0.016 to 0.027 km^3/year . These rates are comparable in magnitude to volumetric increase rates of 0.01

to 0.028 km^3/year , estimated to explain uplift of the caldera from 1923 to 1984 (18).

The line-of-sight displacement inferred with interferometry is consistent with the vertical displacement measured in leveling surveys (19). For the modeled deformation (Fig. 3), the line-of-sight displacement is dominantly vertical motion (20), justifying the direct comparison between InSAR and the leveling surveys (Fig. 4). The reason for the mismatch between the two deformation measurements on the 20 km of the survey line closest to Mount Washburn (MW) (Fig. 4A) is unknown. Because the time intervals over which the measurements were made are different, the mismatch may indicate that the benchmarks near MW subsided another 5 to 10 mm in the 3 months between the June 1993 acquisition date of the slave image in Fig. 2B and the September 1993 leveling survey. Alternatively, either sensitivity of InSAR measurements to atmospheric disturbances (21) or accumulated leveling error may explain the mismatch. The InSAR measurements show a maximum of ~ 15 mm of uplift of the leveling line from July 1995 to June 1997 (Fig. 4C). Although the inferred uplift has not yet been verified by leveling, we do not think

Fig. 2. Six interferograms constructed with the two-pass method (12, 25). The range of colors from violet to red, shown in the color bar at the top of (B), corresponds to one cycle of phase from 0 to 2π (one fringe), representing ~ 28 mm of displacement between a point on the ground and the satellite. The time interval spanned by each interferogram is referenced to the horizontal time axis, where the beginning of each year is marked and labeled. The red, green, and blue bars attached to each interferogram indicate the corresponding time interval. The first three interferograms are sequential, whereas the last three are cumulative over longer time spans (with some overlap). Where coherence is adequate, the interferograms show that deformation extends only slightly beyond the northeast caldera boundary to the northwest-southeast-trending faults mapped by Christiansen (Fig. 1) (2). (A) August 1992 to June 1993. This image shows over 30 mm of inferred subsidence centered in the northeast half of the caldera and closely associated with the SC dome. (B) June 1993 to August 1995. In this image, the center of deformation has shifted to the southwest half of the caldera with over 40 mm of subsidence associated with the ML dome. (C) August 1995 to



September 1996. The main deformation mode is now uplift (~ 20 mm) associated with the SC dome [note the reversed color sequence toward the center of the SC dome, relative to (A) and (B)]. The ML dome still appears to be subsiding slightly during this time interval. (D) August 1992 to June 1995. This image shows ~ 60 mm of subsidence. (E) July 1992 to August 1995, also showing ~ 60 mm of subsidence. (F) July 1995 to June 1997. This image shows over 30 mm of uplift.

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that it is an atmospheric artifact because we see it in two different interferograms (Fig. 2, C and F). Figure 2, C and F, show unique pairings of master/slave scenes (Table 1), implying different atmospheric conditions.

The migration of uplift from the SC to the ML dome suggests that the driving pressure source may lie beneath the SC dome. If this source supplies fluid in a near-vertical pipe-like body, we may use our value of inflation rate and estimates of the driving pressure gradient and fluid viscosity to estimate the

equivalent dimension of such a pipe (22). The fluid may be either magmatic or hydrothermal fluids (23). If the injected magma were basalt, the pipe radius would be about 0.14 m. Rhyolitic magma would require a pipe with a radius of about 14 m, whereas water or steam injection could occur in a pipe with a characteristic dimension of a few millimeters.

We suggest that interacting fluid reservoirs exist beneath the SC and ML domes, with the inflation of each being regulated by flow through two conduits, one beneath SC (conduit

1) and the other connecting SC and ML (conduit 2). The opening of each conduit requires a critical pressure threshold (P_c), perhaps the pressure (P) required to overcome the lithostatic load that would tend to close the conduit walls. Inflation begins beneath SC when the influx of fluids from depth causes P to exceed P_c at conduit 1, with surface uplift occurring at the SC dome. Fluid flow increases pressure in the SC reservoir until it reaches P_c at conduit 2, initiating flow into the ML reservoir and surface uplift above it. Because $P = P_c$ in the SC reservoir, uplift ceases there while the ML reservoir inflates. When P reaches P_c in both reservoirs, and if flow through conduit 1 continues, both should continue inflating. However, if the fluid supply rate through conduit 1 decreased, inflation and concomitant uplift would decline accordingly.

The direction of subsidence migration across the caldera is not known (24), and an outlet during subsidence at conduit 1 would require that the density contrast between fluid and wallrock change sign (22). So we suggest a third conduit as an outlet from the caldera, feeding the shallow hydrothermal features in the northwest corner of the park. In two interferograms (Fig. 2, B and F), the fringes cross the northwest caldera rim, suggesting a connection between the caldera system and the northwest corner of the park. Moreover, both the change from uplift to subsidence in 1985 and the change from subsidence to uplift in 1995 were either initiated or accompanied by the two largest earthquake swarms ever recorded in the park (since 1972, in the USGS–University of Utah earthquake catalog). The two swarms were located just outside of the northwest caldera rim.

Fig. 3. Results of modeling displacement in the interferogram of Fig. 2B with two tabular sill-like bodies embedded in a homogeneous, isotropic, elastic half-space at a depth of 8 km. The fringe interval is the same as shown in Fig. 2B. (A) Synthetic interferogram formed with the best-fitting minimum variance model of Fig. 2B.

The outlines of two tabular bodies that make up the model are shown by white dashed lines. The optimum model was found using code written by Feigl and Dupré (26) to forward model and then apply constrained Monte Carlo searches to estimate best-fitting model parameters (16). (B) Residual interferogram formed by subtracting the model interferogram of (A) from the data (Fig. 2B).

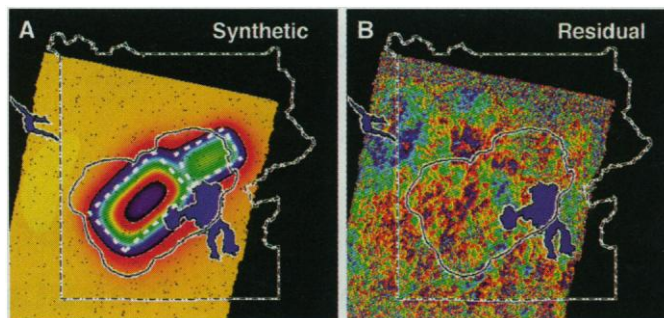
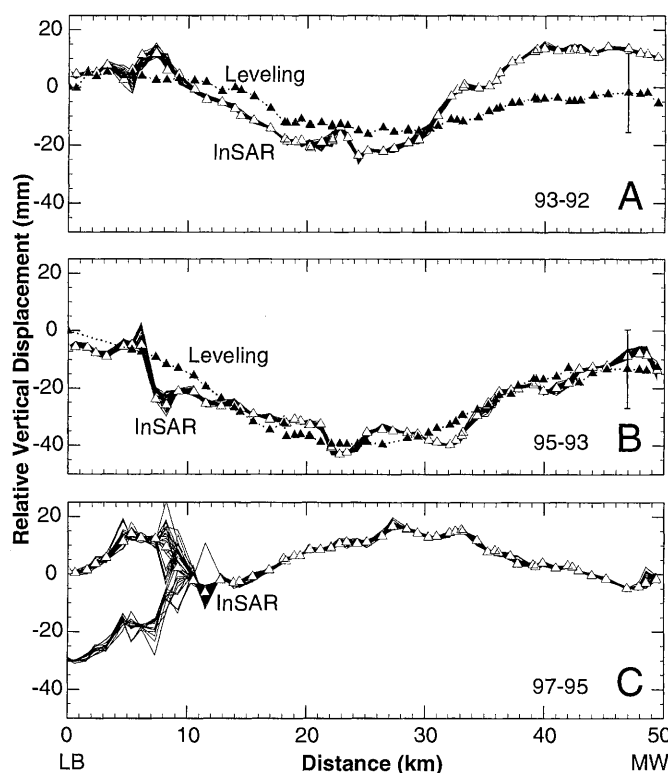


Fig. 4. Vertical displacement inferred from interferometry data at benchmarks on the leveling line shown in Fig. 1. Distance on the horizontal axis is measured along the leveling line in the direction from LB to MW (Fig. 1). Displacement inferred from interferometry (19, 20) is marked with open triangles connected with solid lines, labeled "InSAR." Displacement measured in leveling surveys is marked with solid triangles connected with a dashed line, labeled "Leveling." The vertical error bars on the right side of (A) and (B) show the maximum estimated error in the leveling measurement (~14 mm). The error decreases to ~0 mm at the beginning of the profile. (A) Comparison of vertical displacement inferred from the interferogram in Fig. 2A (August 1992 to June 1993) (19) to that measured from September 1992 to September 1993 at benchmarks (9). (B) Comparison of vertical displacement inferred from the interferogram in Fig. 2B (June 1993 to August 1995) to that measured from September 1993 to September 1995 at benchmarks (9, 10). (C) Vertical displacement inferred at benchmarks using interferogram in Fig. 2F (July 1995 to June 1997) (19). Leveling data after September 1995 are not yet available.



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- inspection. We used inspection to set bounds for a grid search, then accepted the plane that gave a minimum variance-corrected interferogram. Because of poor coherence, the planes used to correct images in Fig. 2, E and F, were more qualitative in nature (there were no minimum variance solutions). The magnitude of tilt across each image is as follows: 30 mm (Fig. 2A), 300 mm (Fig. 2B), 220 mm (Fig. 2C), 300 mm (Fig. 2D), 220 mm (Fig. 2E), and 110 mm (Fig. 2F). These magnitudes are much larger than any signature that could be expected from regional tilting.
14. We assessed the DEM inaccuracies by forming interferograms from SAR images paired from summers of the same year, where the pairs had low altitudes of ambiguity (~25 m). These pairs were sensitive to DEM inaccuracies and (because of the short time interval) fairly coherent. Inaccuracies appear to be generally less than 10 m, with the exception of a small area off the southeast tip of Hebgen Lake (HL) that is in error by ~25 m. This DEM artifact has the largest effect in Fig. 2B (Table 1), where it appears as an area with half a fringe of signal, which could be interpreted as ~14 mm of line-of-sight subsidence. The effect of the DEM artifact near HL is smallest in Fig. 2A (Table 1), where it would appear as an area with only ~1.5 mm of subsidence.
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 20. In comparing leveling and InSAR displacements, we are assuming that the satellite line-of-sight measurements in Fig. 4 represent predominantly vertical deformation. However, the satellite line of sight is inclined 23° to the vertical, so horizontal displacement may be mapped into range displacement. We assessed the magnitude of the error introduced by using this assumption by computing the effect of horizontal displacements in the synthetic (model) interferogram (Fig. 3B). The results suggest that the largest error along the leveling route

introduced by this assumption is less than 3 mm, located where the route crosses the edges of the smaller sill-like body (Figs. 1 and 2).

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22. The volume flux rate (V) of laminar Poiseuille flow in a cylindrical pipe is given by

$$v = \frac{\pi(\rho_s - \rho_l)g}{8\mu} R^4$$

where μ is fluid viscosity, ρ_s is the density of the solid wallrock, ρ_l is the fluid density, g is the acceleration of gravity, and R is the equivalent pipe radius [D. L. Turcotte and G. Schubert, *Geodynamics: Application of Continuum Physics to Geological Problems* (Wiley, New York, 1982), p. 240]. Our volume flux rate is about $0.016 \text{ km}^3/\text{year}$; the driving buoyant pressure gradient $(\rho_s - \rho_l)g$ equals 2000 Pa/m ; and μ is $10^8 \text{ Pa} \cdot \text{s}$ (rhyolite), $1 \text{ Pa} \cdot \text{s}$ (basalt), and 10^{-4} to $10^{-6} \text{ Pa} \cdot \text{s}$ (water and vapor, respectively).

23. We use the term magmatic fluids to refer to magma (molten rock) or volatile gases and liquids released from magma, whereas we use the term hydrothermal fluids

to refer to steam and hot water that are mostly of meteoric origin.

24. From the snapshots of subsidence (Fig. 2) we cannot discern whether the subsidence deformation front moves from southwest to northeast or northeast to southwest. Because the inception of uplift was captured in Fig. 2C, however, we can conclude that the uplift deformation front moves from the northeast to the southwest (Fig. 2F).
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27. We thank our colleagues R. L. Christiansen and J. C. Savage and two anonymous reviewers for thorough technical reviews of the manuscript; R. O. Fournier, D. J. Stevenson, E. Brodsky, D. Massonnet, and N. Pourthie for helpful discussion; and P. Wessel and W. Smith for GMT software [*Eos* **72**, 441 (1991)]. This research was supported by the Volcano Hazards Program, USGS. W.T. was supported by a visiting professorship at the Division of Geological Sciences, California Institute of Technology, while some of the work reported here was carried out.

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Induced Massive Star Formation in the Trifid Nebula?

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The Trifid nebula is a young (10^5 years) galactic HII region where several proto-stellar sources have been detected with the infrared space observatory. The sources are massive (17 to 60 solar masses) and are associated with molecular gas condensations at the edges or inside the nebula. They appear to be in an early evolutionary stage and may represent the most recent generation of stars in the Trifid. These sources range from dense, apparently still inactive cores to more evolved sources, undergoing violent mass ejection episodes, including a source that powers an optical jet. These observations suggest that the protostellar sources may have evolved by induced star formation in the Trifid nebula.

Stars rarely form in isolated events. Observations suggest that, in star-forming regions, once some stars are formed the process spreads further into the parent molecular cloud. The pos-

sibility that stars could trigger the birth of the next generation of stars has long been investigated (1, 2), and HII regions were considered to be good places to look for induced star formation. HII regions form from the photoionization of the molecular gas surrounding the massive newly born stars. The densest condensations of the clouds, cometary globules (CGs), are left behind the ionization front where they are exposed to stellar radiation. Evidence that CGs are sites of star formation is well established based on the presence of infrared (IR) point sources or bipolar outflows (3, 4). Although models suggest that the gravitational collapse should occur in the early stages of photoionization (5), most of the sources found to be associated with CGs display properties typical of relatively evolved objects that have already accreted the bulk of their mass (6). Information about their formation mechanism (spontaneous or radiation-driven) is thus difficult to retrieve, possibly because most studies were done on well-developed HII regions.

The Trifid nebula is a nearby galactic HII region with an estimated distance of 1.68 kpc

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