

X-RAY CRYSTALLOGRAPHY

Beam Splitter Teases Out Phase Secrets

X-ray crystallography, the workhorse of structural biology, may soon be hefting a lighter load. The technique is widely used to produce three-dimensional maps of the jumble of atoms that make up proteins and other molecules. Although the technique is enormously powerful—these days it can decipher the precise structure of molecules containing tens of thousands of atoms—it is also immensely laborious. Mapping out a large molecule often requires researchers to add heavy metals to some crystals of the molecule, scatter x-rays off each crystal in many different orientations, and then compare all of the resulting diffraction patterns. The process can take years of work. But now a Cornell University researcher has come up with a new technique that should considerably speed up the process.

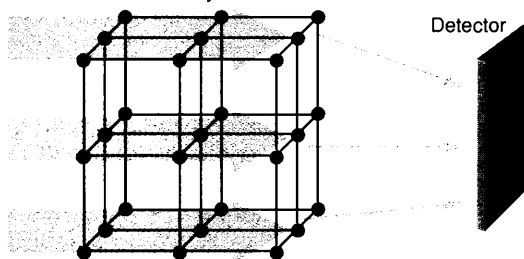
Qun Shen, an x-ray physicist at the Cornell High Energy Synchrotron Source in Ithaca, New York, reports in the 13 April *Physical Review Letters* (PRL) that he has developed a new technique that may do away with the need to make heavy-metal versions of the crystal. Shen uses the crystal itself to split the x-ray beam in two, which results in two beams passing through the crystal at slightly different angles. This produces diffraction patterns that provide the data needed to calculate the molecule's structure.

"There is a lot of potential here," says Michael Rossman, an x-ray crystallographer at Purdue University in West Lafayette, Indiana. He adds that the new technique still has a few bugs to work out before it can be widely adopted. Most important, it requires painstaking work to align the crystal in the first place. But if that procedure can be automated, Rossman and others say the technique could substantially speed the work of protein mappers.

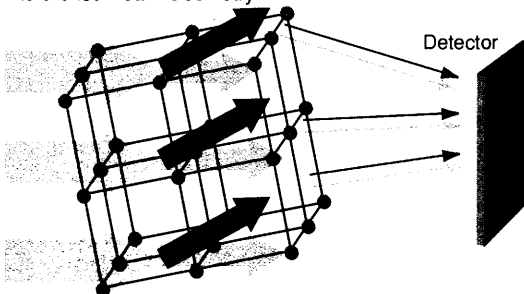
The process of diffraction is the key to x-ray crystallography: X-rays ricochet in slightly different directions from different types of atoms depending on the number of electrons they carry. In a conventional setup, x-rays scatter off atoms in the crystal and hit a detector in a characteristic pattern of dots arrayed in concentric ellipses. This diffraction pattern is the result of interference between ricocheting x-ray waves. Where the peaks and troughs of separate waves rise and fall in unison, they add together, producing a bright spot on the detector. Where the peaks and troughs are out of synch, they cancel each other out, leaving the detector blank. The intensity of the spots and the space between them reveal essential information about the distances between atoms in the crystal. But to know exactly where each atom

sits relative to its neighbors, researchers must find out the relative positions of the scattered x-rays, known as their phase, which is not recorded by the detectors.

Direct-Beam Geometry



Reference-Beam Geometry



Split vision. Tilting the crystal (bottom) splits the x-ray beam and generates diffraction patterns containing phase information.

One of the most common ways researchers get this information is by making a version of the protein crystals with heavy-metal atoms inserted in the core of each protein. The large number of electrons around these metals deflect x-rays moving through the crystal in a different manner. And by com-

paring patterns from crystals with and without the metal, researchers can triangulate the position of the metal atom and then work out the positions of other atoms in the structure relative to that point.

Shen's setup extracts the phase information in a very different way. First, he aligns the crystal precisely so that initially the x-ray beam hits the crystal planes at a very narrow angle. Each plane acts like a mirror, reflecting a large portion of the x-rays in one direction, and so producing two beams passing through the crystal at different angles. Both beams are diffracted by the atoms of the crystal, and as they emerge they interfere with each other, creating a diffraction pattern. The key is that this pattern is now sensitive to phase: Slightly changing the tilt of the crystal changes the angle—and hence phase—of the reflected second beam, which in turn alters the intensity of each spot in the final diffraction pattern. "By looking at the interference between the two beams, we can get the relative phase information of the diffracted waves," says Shen. That information can then be plugged into a software program to solve the protein's structure.

In the PRL paper, Shen reports using this technique to get the phase information needed to solve the structure of semiconductor materials, which are simple to interpret because they have only a small number of atoms in each unit cell. But he and his Cornell colleagues are now trying the technique on a crystallized protein known as lysozyme. If the experiment succeeds, it could help molecular cartographers draw their maps with ease and speed.

—Robert F. Service

PALEOCLIMATOLOGY

A Dusty Ice Age Trigger Looks Too Weak

A study of interplanetary dust has dealt a setback to an upstart proposal about what drives the ice ages. Most climate researchers believe that a slow redistribution of the sun's heat across the globe, driven by cyclic variations in the shape of Earth's orbit, drives the advance and retreat of ice sheets every 100,000 years. But physicist Richard Muller of the University of California, Berkeley, has argued that another kind of orbital cycle, the changing tilt of the orbital plane, could dip Earth in and out of a cloud of interplanetary dust that could somehow alter climate (*Science*, 11 July 1997, p. 183).

Now, on page 874 of this issue, two planetary scientists report calculations suggesting that dust should indeed fall to Earth in a 100,000-year cycle that matches the ice ages.

What drives the cycle, however, is the changing ellipticity of Earth's orbit—the orbital change commonly thought to drive the ice ages—rather than Muller's tilting. And the relatively tiny amount of dust delivered to Earth varies by only a factor of 2 to 3. That's hardly enough to trigger anything so dramatic as an ice age, says geochemist Kenneth Farley of the California Institute of Technology in Pasadena: "It's awfully small to do much of anything."

The tenuous dust cloud that envelops the inner planets is visible to the naked eye on a clear night, but to calculate how much dust may fall to Earth, the orbital behavior of individual dust particles must be known. So Stephen Kortenkamp of the Carnegie Institution of Washington's Department

NAS Elects New Members

The National Academy of Sciences last week announced the election of 60 new members—eight women and 52 men—and 15 foreign associates. The election brings the total number of current active members to 1798 and the number of foreign associates to 310. Following are the newly elected members and their affiliations at the time of election. (For more details, see www2.nas.edu/whatsnew/289e.html)

David E. Aspnès, North Carolina State University; **Bruce J. Berne**, Columbia University; **William A. Brock**, University of Wisconsin, Madison; **A. Welford Castleman**, Pennsylvania State University, University Park; **William L. Chameides**, Georgia Institute of Technology; **Ronald R. Coifman**, Yale University; **Douglas L. Coleman**, Jackson Laboratory, Bar Harbor, Maine; **Elizabeth Anne Craig**, University of Wisconsin, Madison; **Roy G. D'Andrade**, University of California (UC), San Diego; **Ingrid Daubechies**, Princeton University; **Charles A. Dinarello**, University of Colorado School of Medicine; **David L. Donoho**, Stanford University; **William F. Dove**, University of Wisconsin, Madison; **Robert N. Eisenman**, Fred Hutchinson Cancer Research Center, Seattle; **Morris P. Fiorina Jr.**, Harvard University; **E. Norval Fortson**, University of Washington, Seattle; **Perry A. Frey**, University of Wisconsin, Madison; **Susan Gottesman**, National Cancer Institute; **Norma Graham**, Columbia University; **Charles G. Gross**, Princeton University.

Donald A. Gurnett, University of Iowa; **John M. Hayes**, Woods Hole Oceanographic Institution; **Roman Jackiw**, Massachusetts Institute of Technology (MIT); **Thomas H. Jordan**, MIT; **Robert P. Kirshner**, Harvard University; **Miles V. Klein**, University of Illinois, Urbana-Champaign; **Michael S. Levine**, UC Berkeley; **Malcolm A. Martin**, National Institute of Allergy and Infectious Diseases; **Douglas S. Massey**, University of Pennsylvania, Philadelphia; **John J. Mekalanos**, Harvard Medical School (HMS); **James K. Mitchell**, Virginia Polytechnic Institute and State University, Blacksburg; **James M. Moran**, Harvard-Smithsonian Center for Astrophysics; **Craig Morris**, American Museum of Natural History; **Eva J. Neer**, Brigham and Women's Hospital, HMS; **Carl O. Pabo**, Howard Hughes Medical Institute (HHMI) and MIT; **Paul H. Rabinowitz**, University of Wisconsin, Madison.

Sherwin Rosen, University of Chicago; **Joan V. Ruderman**, HMS; **Martin Saunders**, Yale University; **J. William Schopf**, UC Los Angeles; **William R. Schowalter**, University of Illinois, Urbana-Champaign;

Lu Jeu Sham, UC San Diego; **Brian J. Staskawicz**, UC Berkeley; **Paul J. Steinhardt**, University of Pennsylvania; **Melvin E. Stern**, Florida State University; **Audrey Stevens**, Oak Ridge National Laboratory; **Kenneth N. Stevens**, MIT; **Nobuo Suga**, Washington University, St. Louis; **Jack W. Szostak**, Massachusetts General Hospital, HMS; **Lewis G. Tilney**, University of Pennsylvania; **Roger Y. Tsien**, HHMI and UC San Diego; **David B. Wake**, UC Berkeley; **David C. Ward**, Yale University School of Medicine; **Robert G. Webster**, St. Jude Children's Research Hospital, Memphis; **Susan R. Wessler**, University of Georgia, Athens; **Michael S. Witherell**, UC Santa Barbara; **Richard L. Witter**, Michigan State University; **H. Boyd Woodruff**, Soil Microbiology Associates Inc., Watchung, NJ; **Andrew C. Yao**, Princeton University; **Jan A.D. Zeevaert**, Michigan State University.

Newly elected foreign associates, their affiliations, and country of citizenship:

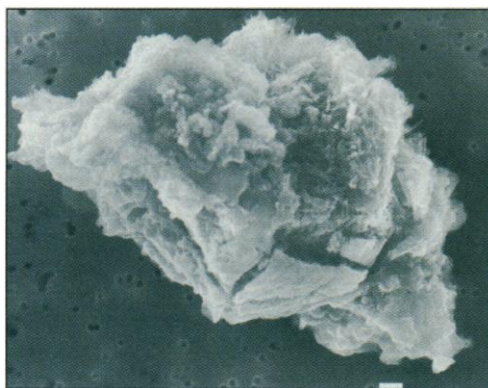
Duilio Arigoni, Swiss Federal Institute of Technology, Zurich (Switzerland); **Peter Doherty**, St. Jude Children's Research Hospital (Australia); **Bryan D. Harrison**, University of Dundee, and Scottish Crop Research Institute, Dundee, Scotland (U.K.); **Richard Henderson**, Medical Research Council Laboratory of Molecular Biology, Cambridge (U.K.); **Hans R. Herren**, International Centre of Insect Physiology and Ecology, Nairobi, Kenya (Switzerland); **Bert Hölldobler**, University of Würzburg (Germany); **Anthony R. Hunter**, Salk Institute for Biological Studies, La Jolla, CA (U.K.); **Edward Irving**, Geological Survey of Canada, Sidney, B.C. (Canada); **Kiyosi Ito**, Kyoto University (Japan); **Maarten Koornneef**, Wageningen Agricultural University, Netherlands (Netherlands); **Olli V. Lounasmaa**, Helsinki University of Technology, Espoo (Finland); **Lelio Orci**, University of Geneva Medical School (Italy); **Roger Penrose**, Oxford University (U.K.); **Romuald Schild**, Polish Academy of Sciences, Warsaw (Poland); **Yasuo Tanaka**, Max Planck Institute for Extraterrestrial Physics, Garching, Germany (Japan).

of Terrestrial Magnetism and Stanley Dermott of the University of Florida, Gainesville, set out to calculate dust orbits by simulating how planetary gravity, sunlight, and the solar wind disperse dust as it is generated by collisions among asteroids. Then they simulated how Earth gravitationally sweeps up the dust as its orbit takes it through the cloud.

The researchers found that the dust cloud spreads too far above and below the plane of the solar system for the tilting of Earth's orbit by a few degrees—Muller's mechanism—to change the amount of dust that reaches Earth. But changing the orbital shape could make a difference. When the orbit is more round, Earth moves more slowly through the cloud and its gravity can more easily pull in dust particles. When the ice ages were at their zenith, so was the amount of atmospheric dust, according to their calculations.

But even then the flux of dust would have been too low to chill the planet, say most observers. Dust that hits the upper atmosphere could affect climate, says atmospheric physicist Donald Hunten of the University of Arizona in Tucson, by vaporizing and recondensing into

tiny particles that might alter the amount of sunlight Earth reflects. But Hunten has calculated that the current dust influx produces a negligible amount of these light-reflecting particles—"[five] orders of magnitude too small to



Speck from an asteroid. Earth's gravity swept up this interplanetary dust particle.

be detected by any means we could imagine," he says, "let alone [great enough] to have an effect on climate." The two- or threefold increase that Kortenamp and Dermott calculate for the height of the ice ages wouldn't make much difference, he adds.

Muller isn't ready to give up on his tilt mechanism, saying that although Kortenamp and Dermott's work "is a great start in doing these sorts of calculations, there's still a lot of room for uncertainty." He notes that Farley's studies of the amount of extraterrestrial dust actually preserved in marine sediments also reveal a 100,000 year cycle—but the timing is the opposite of Kortenamp and Dermott's calculated cycles. "We don't yet understand the accretion of dust onto the Earth," he says.

Kortenamp and Dermott agree that there's room for more dusty analysis. But the only way they see to boost the dust flux to Earth by orders of magnitude would be a catastrophic collision that blows a sizable asteroid to smithereens in the asteroid belt. The resulting pulse of dust could be four orders of magnitude higher than normal for up to a million years, they say, and might account for several intervals of increased dust flux in Farley's 70-million-year sedimentary record. The resulting climate change might even have triggered gradual extinctions through climate change, they say. But that's another speculation begging for a test.

—Richard A. Kerr

SOURCE: J. P. BRADLEY AND D. BROWNLEE