

# Reports

## Volcanic Gases: Hydrogen Burning at Kilauea Volcano, Hawaii

**Abstract.** *Spectroscopic evidence for hydrogen burning in air was obtained at Kilauea Volcano. The abundance of hydrogen required to support combustion is consistent with that predicted for gases in equilibrium with typical Hawaiian tholeiitic basalt.*

Volcanic gas is usually collected under low-temperature fumarolic conditions, and, as a result, the composition of the gas may have been altered by reaction with subsurface rocks and perhaps groundwater. Only on rare occasions, and then with some difficulty, can collections be made at high temperatures where gas and molten lava are in intimate contact. The gases collected under these conditions are in thermodynamic equilibrium (1), but in the special case of gases burning at a vent thermodynamic equilibrium with the ambient air is achieved after combustion. In Hawaii flames are occasionally seen from burning volcanic gas, usually during or immediately after an eruption, and can, in certain exceptional cases such as that reported here, be studied by spectroscopic methods.

Previous attempts at flame spectroscopy at several volcanoes (2-4) have revealed emission bands of CuCl, lines of Na and K, and some unidentified continua, but have given little information on the combustion process occurring in the flames. We report here the identification of spectral features observed at Kilauea Volcano that establish the flames as originating from the combustion of  $H_2$  in air. Although the burning of  $H_2$  had long ago been suggested as the cause of certain volcanic flames (4), reliable evidence for  $H_2$  burning was previously unavailable and the possibility remained that other combustibles, such as CO, were the prime cause of the flames seen at Kilauea and elsewhere.

An 11-hour episode of lava fountaining on the caldera floor of Kilauea Volcano, Hawaii, began at 0855 H.S.T. on 14 August 1971. Considerable portions of the caldera floor were covered with new lava from an echelon line of vents preventing access except to the

southwesternmost extremity of one fissure. Large and colorful flames were seen emanating from these vents after sunset, following the termination of fountaining.

Our observing party approached to within 20 m of a line of flaming vents, and at the edge of a fresh lava flow we set up a small field spectrograph (5) especially constructed for the study of volcanic flames. The line of vents was about 35 m long and consisted of six separate flaming orifices. Flames at one end of the line were greenish-blue, and those at the other end were yellow-orange, with

intermediate colors between. Hour-long exposures, shown in Fig. 1, were made of both the greenish-blue and the yellow-orange flames between 2005 and 2400 H.S.T. (by 2400 H.S.T. their intensities had diminished greatly). At maximum activity the flames rose about 2 m above the rims of the orifices. The most striking differences between the two extremes of flame color are the intensity of the Na doublet (blended because of overexposure) and the faint continuum from about 5800 to 6200 Å.

Between 4900 Å, where the filter begins to transmit, and 5300 Å no radiation was recorded. A weak continuum appears between about 5300 and 5500 Å. This is most likely part of the Na system; flames showing strong resonance lines of alkali metals also exhibit continuous emission (6). The resonance lines of atomic Na at 5890 and 5896 Å are seen as an impurity in nearly all flames. The dark gap in the spectrum near 6200 Å corresponds to a dip in the sensitivity of the Eastman Kodak class F plates used in this study. In the spectral region covered by this series of plates, emission features of  $CO_2$ , due to the combustion of CO in air, occur (6) but are not seen in our records. We originally considered CO combustion

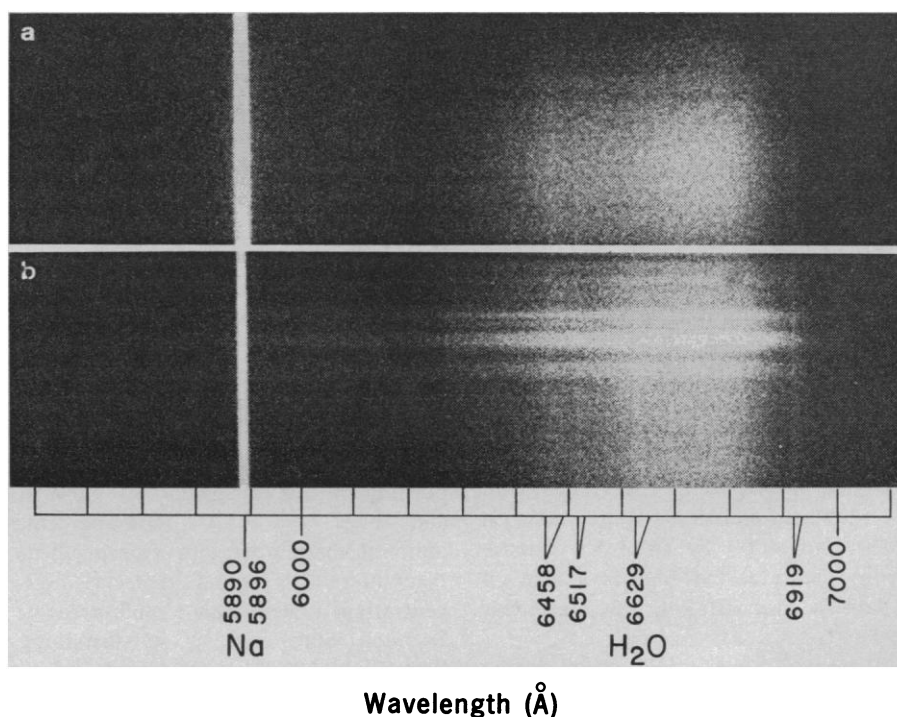
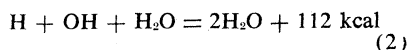


Fig. 1. Spectra of volcanic flames, 14 August 1971. (a) Hour-long exposure of a yellow-orange flame; (b) similar exposure of a greenish-blue flame emanating from a nearby vent. The Na lines are blended by overexposure, but note the relative intensities in the two flames. The four most intense bands of the  $H_2O$  emission spectrum are indicated.

to be the prime candidate for the generation of volcanic flames, owing to the often relatively high concentration of this gas in volcanic emanations. We are unable to establish an upper limit on the concentration of CO in the flames we observed. Bands of S<sub>2</sub>, which should result from combustion of H<sub>2</sub>S in air, are not seen in our records, but in a spectrum of the lava fountain at the 1967–1968 eruption of Halemaumau Cruikshank found marginal evidence for one relatively intense S<sub>2</sub> band (7).

Between about 6450 and 6915 Å there appears a series of bands corresponding to the four most intense members of a system of H<sub>2</sub>O emission bands first discovered in high-pressure premixed H<sub>2</sub>-O<sub>2</sub> flames (8). A reproduction of a laboratory spectrum showing these bands is given by Gaydon [plate 1 in (6)].

In a laboratory study of the H<sub>2</sub>O bands, Kitagawa (8) considered their formation by the reactions



Since the shortest wavelength bands of the observed system require about 50 kcal/mole, these reactions would lead to the necessary excitation if most of the energy goes into the vibration states of the H<sub>2</sub>O molecule. The non-metathetical reaction (Eq. 1) does not normally occur, although it is stoichiometrically correct. In order to determine which of the two reactions occurs in the flames at Kilauea, the emission features of OH were sought. The strongest accessible emission bands (0, 0 and 0, 1) occur in the ultraviolet region between 3064 and 3472 Å (6, p. 79). These bands were invisible in spectra of the Halemaumau lava fountain in 1967 and 1968 (9, 10). We observed flames at another eruption of Kilauea in late September 1971 but could not detect the ultraviolet bands. Delsemme (3) failed to detect OH emission in flames in the Congo, and we found no emission features between 3000 and 4000 Å from an intense yellow flame at the Alae lava lake on Kilauea's east rift on 13 December 1972.

Present evidence, then, suggests that excited H<sub>2</sub>O molecules form without the intermediate formation of the OH radical. Bimolecular branching reactions involving H and O and the many combinations of radicals and stable molecules all have activation energies

too low to yield the observed bands of H<sub>2</sub>O (11). The OH radical, if present, presumably would enter into the flame chemistry by Kitagawa's Eq. 2.

Hydrogen is normally a minor constituent of volcanic gases collected from fuming vents (12). Its concentration can be understood in terms of the equilibrium with O<sub>2</sub>; the O<sub>2</sub> fugacity is established by the oxidation states of iron in a rock melt-volatile system (13). The quartz-fayalite-magnetite buffer system (QFM), defined by Eugster and Wones (14), has provided a reliable approximation of the O<sub>2</sub> fugacity in equilibrium with basaltic melts at temperatures near the liquidus (15). Calculated O<sub>2</sub> fugacities for analyzed volcanic gases (1) are in excellent agreement with those predicted for gases coexisting with the QFM buffer at temperatures above about 1200°K. Electron microprobe analysis of coexisting ilmenite-magnetite in a Hawaiian tholeiite (16) and in situ measurements of O<sub>2</sub> concentrations (17) at 1300°K agree closely with the values predicted for the QFM system. Thus the O<sub>2</sub> fugacity of the gas phase coexisting with Hawaiian tholeiite can be adequately specified for the temperature range from 1300° to 1525°K. If we assume thermodynamic equilibrium at these high temperatures in a typical melt, the knowledge of O<sub>2</sub> fugacity permits computation of the ratio of partial pressures. In the temperature range considered,  $p_{\text{H}_2\text{O}}/p_{\text{H}_2} = 54 \pm 2$  and is relatively insensitive to temperature.

The value of  $p_{\text{H}_2}$  can be readily calculated once  $p_{\text{H}_2\text{O}}$  in the volcanic gas is determined. From infrared absorption studies, Naughton *et al.* (10) assigned a value of 95 percent for the H<sub>2</sub>O content of Hawaiian volcanic fume, a value which is supported by analyses of collected samples. At 1475°K,  $p_{\text{H}_2}$  is 14 torr (1.8 mole percent). This should be considered an upper limit because various processes tend to deplete H<sub>2</sub> selectively (18).

The lower flammability limit of H<sub>2</sub> burning in air is given by Coward and Jones (19) as 4.0 mole percent, although the temperature is unspecified. Naughton (20) found that, at concentrations of less than 2 mole percent, H<sub>2</sub> will burn with a self-sustaining flame in air if the orifice temperature is greater than 975°K, whereas gas containing less than 0.5 mole percent H<sub>2</sub> can ignite if the orifice temperature is about 1125°K. In addition to the hot orifice walls, condensed solids in the fume may provide the source of heat

for ignition. Large flames that immediately follow the cessation of lava fountaining from fissures at the summit and on the flanks of the mountain normally persist for several hours after fountaining and then slowly diminish and disappear. Their disappearance is probably related to the cooling of the orifice below the flammability limit for the molar fraction of H<sub>2</sub> at that particular eruptive event, rather than to a change of gas composition.

DALE P. CRUIKSHANK

DAVID MORRISON

*Institute of Astronomy, University of Hawaii, Honolulu 96822*

KENNETH LENNON

*Department of Chemistry, University of Hawaii*

#### References and Notes

1. A. J. Ellis, *Amer. J. Sci.* **255**, 416 (1957); E. F. Heald, J. J. Naughton, I. L. Barnes, *J. Geophys. Res.* **68**, 545 (1963); B. Nordlie, *Amer. J. Sci.* **271**, 417 (1971).
2. J. Verhoogen, *Amer. J. Sci.* **237**, 656 (1939); K. J. Murata, *ibid.* **258**, 769 (1960); N. A. Kozyrev, *Izv. Gl. Astron. Observ. Pulkovo* **24** (4), 76 (1966).
3. A. H. Delsemme, *Bull. Seances Acad. Roy. Sci. Outre-Mer* **6** (new ser.) fasc. 3, 507 (1960).
4. Flames at Kilauea were studied with spectroscopes long ago. In 1893, W. Libbey, Jr., used a visual spectroscope at the edge of the Halemauau lava lake to study pale greenish flames that accompanied explosions of lava from the lake surface. He "failed to get much information concerning them, aside from the fact that they seemed the exact counterpart of the hydrogen flame from a Bunsen burner" [*Amer. J. Sci.* **47** (ser. 3), 371 (1894)]. In 1883, the French physicist J. Janssen visited Kilauea briefly and used a spectroscope to study the flames at Halemauau. He reported [*C. R. Acad. Sci.* **97**, 601 (1883)] the detection of Na, H<sub>2</sub>, and hydrocarbons in the flames. Janssen's report is curious in that he does not specify the criteria for his identification of H<sub>2</sub>. The Balmer lines, which lie in the visible portion of the spectrum and in 1855 had been attributed to H<sub>2</sub>, are not excited at the low temperature of volcanic flames. In 1880, G. D. Liveing and J. Dewar [*Proc. Roy. Soc. London* **30**, 494 (1880)] and W. Huggins (*ibid.*, p. 576) had reported the "ultraviolet water vapour bands" from H<sub>2</sub> burning in air, later shown to be due to the free OH radical, but these bands are invisible to the eye with a spectroscope. Liveing and Dewar (*ibid.*) found the flames of H<sub>2</sub> burning in O<sub>2</sub> luminous and continuous but showing no structure directly attributable to the spectrum of hydrogen. T. A. Jaggar [*Amer. J. Sci.* **44**, 161 (1917)] reported detailed field observations of flames at Kilauea, speculating that they originate from H<sub>2</sub> burning in air.
5. The spectrograph uses a plane grating (600 lines per millimeter, blazed for 12,000 Å, first order) and is used in the first and second orders. Dispersion in the second order is 100 Å/mm. This project began in 1967 under the sponsorship of G. P. Kuiper and was supported through 1968 by a NASA institutional grant to the University of Arizona.
6. A. G. Gaydon, *The Spectroscopy of Flames* (Chapman & Hall, London, 1957).
7. D. P. Cruikshank, unpublished data.
8. T. Kitagawa, *Rev. Phys. Chem. Jap.* **13** (No. 2), 96 (1939).
9. D. P. Cruikshank, in preparation.
10. J. J. Naughton, J. V. Derby, R. B. Glover, *J. Geophys. Res.* **74**, 3273 (1969).
11. J. N. Bradley, *Flame and Combustion Phenomena* (Methuen, London, 1969), p. 74.
12. D. E. White and G. A. Waring, *U.S. Geol. Surv. Prof. Pap.* **440-K** (1963); J. B. Finlay-

- son, I. L. Barnes, J. J. Naughton, in *The Crust and Upper Mantle of the Pacific Area*, L. Knopoff, C. Drake, P. Hart, Eds. (Monograph 12, American Geophysical Union, Washington, D.C., 1968).
13. G. Kennedy, *Amer. J. Sci.* **246**, 529 (1948).
  14. H. Eugster and D. Wones, *J. Petrology* **3**, 82 (1962).
  15. I. Carmichael and J. Nicholls, *J. Geophys. Res.* **72**, 4665 (1967).
  16. T. Wright and P. Weiblen, *Geol. Soc. Amer. Spec. Pap. No. 115* (1967), p. 242.
  17. M. Sato and T. L. Wright, *Science* **153**, 1103 (1966).
  18. Once the gas phase has been exsolved from the melt, the  $O_2$  and  $H_2$  pressures are no longer buffered. Any process selectively removing  $H_2$  will induce a degree of disequilibrium, which will be followed by the dissociation of  $H_2O$  vapor. (This will restore the  $H_2$  to approximately its previous pressure but will greatly increase the  $O_2$  pressure.) Two possible processes are the introduction of small amounts of atmospheric  $O_2$  and the diffusion of  $H_2$  through solid lava crusts, the latter having been postulated by Sato and Wright (17) to explain the oxidation of solidified basaltic lava.
  19. H. F. Coward and G. W. Jones, *U.S. Bur. Mines Bull.* **503** (1952).
  20. J. J. Naughton, personal communication.
  21. We thank Dr. D. W. Peterson, scientist-in-charge of the Hawaiian Volcano Observatory, U.S. Geological Survey, and his staff, particularly Dr. D. A. Swanson. We thank the superintendent of the Hawaiian Volcanoes National Park for permission to conduct the spectroscopic project at Kilauea. Research supported under funds from the National Geographic Society.

5 October 1972; revised 9 April 1973

## Ground-Based Thermal Infrared Surveys as an Aid in Predicting Volcanic Eruptions in the Cascade Range

**Abstract.** Detailed, ground-based thermal surveys of selected slopes of Mt. Rainier and Mt. Lassen have provided quantitative data for the construction of thermal contour maps of portions of the flanks of these volcanoes. Preliminary surveys substantiate the existence of one suspected thermal anomaly on Mt. Rainier.

Recent eruptions of supposedly extinct volcanoes such as Mt. Arenal in Costa Rica in 1968 and the recognition of large, geologically recent, lahars (mudflows from a volcano) in volcanic areas have led to an increase in the monitoring of potentially hazardous volcanoes. Volcanic eruptions often produce destructive quantities of lava and ash. However, lahars may pose even greater destructive threats because of their possible speed and size. Lahars covering tens of square miles have been found, for example, in many of the valleys adjacent to Mt. Rainier, Washington (1). It is possible that many of these lahars formed as a result of volcanic warming of portions of the ice- and debris-covered volcano and were not necessarily related to simultaneous outpourings of volcanic materials.

A variety of techniques are now being used on Cascade Range volcanoes in the western United States to predict volcanic eruptions. These include seismic monitoring, intermittent periodic aerial thermal infrared imagery, tiltmeter studies, geodetic surveys, and ground-based quantitative thermal scanning.

We report here preliminary results of one of the field-monitoring techniques, namely, ground-based quantitative thermal scanning of the slopes of some of the most potentially dangerous Cascade Range volcanoes. This particular thermal technique was first used by Birnie (2) on certain Central Ameri-

can volcanoes. Radiometry was also used by Tazieff and Jatteau to determine the temperatures of gases emitted at Mt. Etna, Sicily (3).

In our thermal technique we utilize a PRT-5 precision radiation thermometer (Barnes Engineering Company) (4). The radiation thermometer measures the radiant exitance between 9.5 and 11.5  $\mu\text{m}$  of any target within its 0.14° (2.44 milliradians) field of view. This particular band pass substantially reduces errors normally induced by atmospheric absorption. The 0.14° field of

view of the instrument makes it possible to integrate the radiant exitance of a circular area 1.0 m in diameter at a distance of 401 m from the sensor. The PRT-5 yields precise readings of apparent surface temperatures ranging from -20° to +75°C. The precision of the instrument is 0.1°C.

The radiation thermometer is set up on prominent view sites around each volcano. The PRT-5 optical scanning head is mounted on a goniometer-equipped tripod capable of reproducible vertical and horizontal movements as small as 15'. Readings are taken in the very early morning before the first rays of the sun strike the mountainside, thus minimizing the differential solar heating of the surface material from the preceding day. The slope is scanned horizontally from the top down, and the temperatures are recorded at 20-minute intervals. No corrections are made for the emissivity of the surface material, the atmospheric absorption of the infrared radiation, or the temperature change due to the adiabatic cooling rate with increasing elevation. The vegetation-free volcanic surfaces are approximately constant in composition; therefore, the emissivity is approximately constant but not equal to unity.

The thermal data were collected on clear mornings in August and September 1972, the season of minimum mountain snow cover. Recorded temperatures were rounded off to the nearest degree Celsius and then they were plotted on photographs of the mountain slopes taken from the monitoring sites, and the isotherm contours were

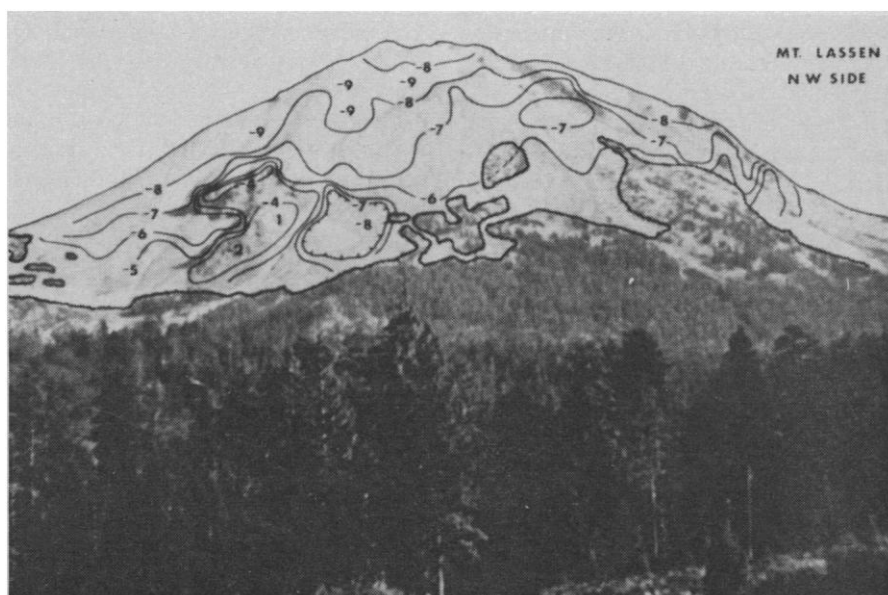


Fig. 1. Northwest side of Mt. Lassen (elevation, 3189 m), California, scanned 24 September 1972, from the north shore of Manzanita Lake (elevation, 1784 m). The distance between the station and the summit is approximately 8.1 km.