further oxidation of NO to NO2 and the rate of ozone formation. This behavior can be explained by a chain reaction consisting of reactions 1, 2, 9, and 10:

$$HO_{2} + NO_{2} \rightarrow HNO_{2} + O_{3} \qquad (9)$$
  
$$HNO_{2} + h_{\nu} \rightarrow OH + NO \qquad (10)$$

This chain reaction converts NO<sub>2</sub> to NO and will compete with reactions 1 through 3 under conditions of large [NO<sub>2</sub>]/[NO].

The carbon monoxide had 0.15 percent  $Fe(CO)_5$  as its principal impurity (the hydrocarbon impurities were negligible, being less than 5 ppm). When iron carbonyl was not removed from the CO, the rate of oxidation of NO to NO<sub>2</sub> was accelerated even more than is shown in Fig. 1. In the presence of unpurified CO, only 60 minutes elapsed before all the NO was converted to  $NO_2$  (defined as when  $[NO] = [O_3]$ ). By comparison, with purified CO, 100 minutes elapsed (see Fig. 1); without CO, 130 minutes elapsed. Since the concentration of  $Fe(CO)_5$  is an order of magnitude smaller than that of the NO, its pronounced effect can be explained only by a chain reaction. Because the carbonyl undergoes decomposition in sunlight (6) and is probably vulnerable to attack by oxygen atoms, a possible chain mechanism for its effect might include the following reactions:

$$Fe(CO)_5 + O \rightarrow Fe(CO)_4 + CO_2$$
 (11)

 $\operatorname{Fe}(\operatorname{CO})_5 + \stackrel{h\nu}{\rightarrow} \operatorname{Fe}(\operatorname{CO})_x + (5-x)\operatorname{CO},$ *x* < 5 (12)

 $Fe(CO)_x + O_2 \rightarrow Fe(CO)_xO_2$ (13)

 $Fe(CO)_{x}O_{2} + (NO \text{ or } CO) \rightarrow$  $Fe(CO)_xO + (NO_2, CO_2)$ (14)

 $Fe(CO)_{x}O + (CO \text{ or } NO) \rightarrow$  $Fe(CO)_x + (CO_2, NO_2)$ (15)

Other reaction mechanisms may be equally possible, but reactions 11 through 15 do show that the observed effects of  $Fe(CO)_5$  can be accounted for in a plausible manner. Iron carbonyl is highly reactive in other freeradical systems: trace amounts of  $Fe(CO)_5$  either quench hydrocarbonair flames or reduce their flame velocities substantially (7); this compound formerly was used as an antiknock agent in gasoline (6).

KARL WESTBERG

NORMAN COHEN Aerospace Corporation,

El Segundo, California 90245 K. W. WILSON

Stanford Research Institute, Irvine, California 92664

12 MARCH 1971

## **References and Notes**

- 1. P. A. Leighton, Photochemistry of Air Pollution (Academic Press, New York, 1961), pp. 146-168, 183, 228.
- 146-168, 183, 228. D. L. Baulch, D. D. Drysdale, A. C. Lloyd, "High Temperature Reaction Rate Data No. 1" (Department of Physical Chemistry, The University, Leeds, England, 1968). J. Heicklen, K. Westberg, N. Cohen, in *Chemical Reactions in Urban Atmospheres*, C. S. Tuesday, Ed. (Elsevier, Amsterdam, in press)
- 3
- press). J. Doyle, Environ. Sci. Technol. 4, 907 4. G. (1970).
- J. Bufalini, ibid. 2, 703 (1968). 6. H. Remy, Treatise on Inorganic Chemistry

(Elsevier, Amsterdam, 1956), vol. 2, p. 289; N. V. Sidgwick, The Chemical Elements and N. V. Slagwick, The Chemical Elements and Their Compounds (Clarendon, Oxford, 1950), vol. 2, p. 1369; G. K. Rollefson and M. Bur-ton, Photochemistry and the Mechanisms of Chemical Reactions (Prentice-Hall, New York, 1939), p. 362.

- 7. G. Lask and H. Gg. Wagner, Symp. Combust. 8, 433 (1962); W. J. Miller, Combust. Flame 13, 210 (1969).
- We thank Prof. J. Heicklen and Drs. A. Bockian and S. W. Benson for valuable suggestions and advice, and I. Sauer and B. Kurtin for technical assistance in the experiments.
- 14 September 1970; revised 14 December 1970

## Sulfur Isotopes in Swaziland System Barites and the Evolution of the Earth's Atmosphere

Abstract. Sedimentary barites from the Swaziland System of South Africa (more than 3000 million years old) have sulfur-34 ratios that are enriched by only 2.5 per mil with respect to contemporary sulfides. To explain this small fractionation, it is proposed that oxygen pressure in the earth's atmosphere was very low and that local oxidation occurred in a photosynthetic layer of the ocean.

In this report we interpret the significance of sulfur isotope data from an isolated occurrence of sulfate separated in time by 2000 million years from any well-understood deposits. We hope to show that the deposit is sedimentary, that its sulfate may reasonably represent contemporary oceanic sulfate, and that its sulfur isotope composition may indicate roughly the oxygen content of the early atmosphere.

Beds of barite alternating with green chert, shale, and pebble conglomerate occur in a 20-m-thick zone of the Fig Tree or Upper Onverwacht Group, Swaziland Sequence, in the Barberton Mountain Land, South Africa. These occurrences are mentioned by Visser (1), and several of them are shown on the 1:50,000 map of the Barberton area published by the South African Geological Survey in 1955; they appear as an 8-km-long band, 11 km southeast of Barberton. The barite is recrystallized, but we consider it to be sedimentary for the following reasons: (i) it occurs in conformable layers, and cross-cutting veins have not been observed; (ii) the zone of barite beds is persistent within the same horizon for several kilometers along strike; (iii) no accompanying sulfide mineralization has been observed; and (iv) the baritechert association found here is common in sedimentary barite deposits.

Strontium isotopic data provide strong evidence that these deposits were formed prior to the main metamorphic event of 3000 million years ago that affected the Barberton Mountain Land.

Whereas the <sup>87</sup>Sr/<sup>86</sup>Sr ratio established in graywackes and shales by this metamorphic event was 0.712 (2), the <sup>87</sup>Sr/<sup>86</sup>Sr ratio in eight Barberton barites ranges from 0.70088 to 0.70172 (3). This lower <sup>87</sup>Sr/<sup>86</sup>Sr ratio precludes derivation of the cations from a source with a high Rb/Sr ratio and suggests that they were derived from the upper mantle during volcanism associated with sedimentation (4).

Table 1 lists  $\delta^{34}$ S values (5) for barites and for sulfides from approxi-

Tabl	e 1.	Values	$\mathbf{of}$	$\delta^{34}S$	for	barites	and	sulfides
of the	ie S	wazilan	d S	Syste	m.			

Sample	$\delta^{34}S$	Remarks
		Barites
A185 9B2 B11 B20	+3.5 +3.1 +3.1 +3.8	Coarse green barite associated with chert, central zone
B14	+3.5	Barite and conglomerate, west workings
B15	+3.4	Elongated barite crystals, east workings
B17	+3.7	Barite in conglomerate, easternmost part of district
Appro:	ximately	contemporaneous sulfides
A177	+1.1	Disseminated pyrite in graywacke, Fig Tree Group
<b>R</b> 5	-0.1	Pyrite in Fig Tree iron formation
R6	+1.6	Pyrite from dolomite bed underlying barite
	-0.9	Galena in quartz vein, Rosetta mine (19)

mately equivalent stratigraphic horizons of the Swaziland Sequence. Sulfur isotope variations are small within each group, and, unlike more recent sulfatesulfide pairs, there is only a 2.6 per mil separation between barite and pyrite. Before presenting an analysis of these data, we shall briefly review the present sulfur cycle.

In the present ocean  $\delta^{34}S$  of sulfate is +20.4 per mil, and the ocean is a reservoir for  $1.3 \times 10^{15}$  metric tons of  $SO_4^{2-}$  (6). Sulfides precipitated from the ocean vary considerably in their isotopic composition, but averages of a large number of determinations show stratified sulfide deposits associated with volcanics to be 17.5 per mil lower in  $\delta^{34}S$  and other sedimentary sulfide deposits to be 11.7 per mil lower in  $\delta^{34}$ S than contemporary sulfates (7). The average  $\delta^{34}$ S for sulfide in basic magmas is probably about +1 per mil (8). The accepted mechanism for enrichment of ocean sulfate in <sup>34</sup>S is the organic reduction of sulfate to sulfide accompanied by a depletion of the sulfide in <sup>34</sup>S by 15 per mil or more (6). Since this is the only quantitatively important reaction in the sulfur cycle that produces a large isotope fractionation, the cumulative effect is an increase in  $\delta^{34}S$  of the sulfate reservoir (6, 9). There is little fractionation between crystalline sulfate and dissolved sulfate; and measurements of Phanerozoic evaporites have given a detailed record of past  $\delta^{34}S$ variation in the ocean, ranging from +11 per mil in the Permian to about +30 per mil in the Cambrian (6). Far fewer data exist for barite than for evaporite minerals, but most syngenetic and oceanic barites have sulfur isotope compositions similar to the sulfate composition of the contemporary ocean (7, 10, 11). In some environments sedimentary barite can have  $\delta^{34}$ S values considerably higher than the  $\delta^{34}$ S of the ocean (11).

Sulfates are formed by other processes, such as low-temperature volcanic reactions. These reactions generally are characterized by sulfate-sulfide isotope fractionation similar in magnitude to fractionation between sedimentary pairs (12). Only supergene sulfate and sulfate from carbonaceous chondrites consistently have low  $\delta^{34}$ S values comparable to the Barberton barites; and, in carbonaceous chondrites, sulfate sulfur is generally slightly depleted in <sup>34</sup>S with respect to sulfide sulfur (13).

We consider the data in Table 1 to be strong evidence that the geochemical cycle of sulfur during deposition of the Swaziland Sequence was grossly different from the present sulfur cycle. A possible starting point in evaluating these data is Holland's model for the evolution of the earth's atmosphere (14). He proposed that at roughly the time of deposition of the Swaziland System, 3200 million years ago,  $H_2S$ was replaced or supplemented in the atmosphere by SO<sub>2</sub>.

We propose a model for deposition of the Swaziland Sequence in which the present large oceanic sulfate reservoir did not exist. Cations for the Barberton barites were provided by volcanic and volcanogenic sources; oxidation occurred in an oxidizing photosynthetic zone near the ocean surface in which, as Cloud has suggested (15), primitive organisms may have used electron donor reactions such as  $Fe(OH)_2$  +  $OH^- \rightarrow Fe(OH)_3 + e^-$  instead of reactions involving free oxygen. Under these conditions, green or purple sulfur bacteria could oxidize  $H_2S$  or  $S^{2-}$  to  $SO_4^{2-}$  in the photosynthetic layer. In the presence of anomalous concentrations of Ba<sup>2+</sup>, precipitation of BaSO<sub>4</sub> would follow. Kaplan and Rittenberg (16) have shown that oxidation of sulfide to  $SO_4^{2-}$  by Chromatium sp. produces little isotopic fractionation. Thus this model is consistent with the very small difference in  $\delta^{34}S$  between barite and sulfides from the Swaziland Sequence. Furthermore, Harrison and Thode (17) have demonstrated that, in the concentration range of  $6 \times 10^{-4}$  to  $1 \times 10^{-5}$  mole/liter of SO<sub>4</sub><sup>2-</sup>, biological sulfate reduction results in fractionation factors of 1.003 to 0.997 (that is, a 3 per mil depletion of  $\delta^{34}$ S in the sulfide at  $6 \times 10^{-4}$  mole/liter and a 3 per mil enrichment at  $1 \times 10^{-5}$  mole/liter of  $SO_4^{2-}$ ), so that it is unlikely that large fractionation could occur as long as the concentration of  $SO_4^2$  in the ocean was controlled by cations like Ba<sup>2+</sup> and Sr<sup>2+</sup>.

Recently Chukhrov et al. (18) have presented evidence for <sup>34</sup>S enrichment in sulfate minerals of rocks metamorphosed at about 2500 million years ago. Their data presumably are obtained from low percentages of sulfurbearing minerals such as lazurite in veins and pegmatites in highly metamorphosed rocks that sometimes contain  $H_2S$  and native sulfur. We suggest that great care should be taken to distinguish depositional and postdepositional processes in these rocks.

Although the Barberton Mountain Land data perhaps can be explained in other ways, we believe they imply that the sulfur cycle was different and that the atmosphere was less oxidizing in the Early Precambrian than at present.

EUGENE C. PERRY, JR. Minnesota Geological Survey, University of Minnesota, Minneapolis 55455

JAN MONSTER

## McMaster University,

Hamilton, Ontario, Canada

THOMAS REIMER

University of the Witwatersrand, Johannesburg, South Africa

## **References and Notes**

- 1. D. J. L. Visser, Special Publication 15 (Union of South Africa Geological Survey, Pretoria, 1956).
- 2. H. L. Allsopp, T. J. Ulrych, L. O. Nicolay-
- Sen, Can. J. Earth Sci. 5, 605 (1968).
   E. C. Perry, Jr., V. R. Murthy, T. Reimer, paper presented at the annual meeting of the Geological Society of America, Milwaukee. 1970.
- 4. For a more detailed discussion of strontium data, see V. R. Murthy and E. C. Perry, Jr., in preparation.

5. 
$$\delta^{34}\mathbf{S} = \left[\frac{(^{34}\mathbf{S}/^{32}\mathbf{S}) \text{ sample}}{(^{34}\mathbf{S}/^{32}\mathbf{S}) \text{ standard}} - 1\right] \times 1000$$

The reference standard is troilite from the Canyon Diablo meteorite. Analyses were per-formed at McMaster University with use of the techniques and equipment described by H. G. Thode and J. Monster [Amer. Ass. Petrol. Geol. Bull. 54, 627 (1970)].

- 6. W. T. Holser and I. R. Kaplan, Chem. Geol. 1, 93 (1966).
- 7. D. F. Sangster, Proc. Geol. Ass. Can. 19. 79 (1968).
- 8. M. Shima, W. H. Gross, H. G. Thode, J. Geophys. Res. 68, 2835 (1963). 9. C. E. Rees, Earth Planet. Sci. Lett. 7, 366
- (1970).10. G. Anger, H. Puchelt, W. Ricke, Econ. Geol.
- G. Anger, H. Pucneit, W. Kicke, Econ. Geot. 61, 511 (1966).
   M. Solomon, T. A. Rafter, M. L. Jensen, Mineral Deposita 4, 172 (1969); E. D. Gold-berg, B. L. K. Somayajulu, J. Galloway, I. R. Kaplan, G. Faure, Geochim. Cosmochim. Astro. 22, 297 (1960). Acta 33, 287 (1969).
- 12. H. Sakai, Geochim. Cosmochim. Acta 12, 150 (1957)
- 13. I. R. Kaplan and J. R. Hulston, ibid. 30, 479 (1966 14. H.
- 479 (1966).
  H. D. Holland, in Petrologic Studies: A Volume in Honor of A. F. Buddington, A. E. J. Engel, H. L. James, B. F. Leonard, Eds. (Geological Society of America, Denver, 1962).
- (Geological Society of America, Denver, 1962).
  15. P. E. Cloud, Science 160, 729 (1968).
  16. I. R. Kaplan and S. C. Rittenberg, J. Gen. Microbiol. 34, 195 (1964).
  17. A. G. Harrison and H. G. Thode, Trans. Faraday Soc. 54, 84 (1958).
  18. F. V. Chukhrov, V. I. Vinogradov, L. P. Ermilova, Mineral Deposita 5, 209 (1970).
  19. W. V. Ault and J. L. Kulp, Geochim. Cos-mochim Acta 16 201 (1959).

- mochim. Acta 16, 201 (1959). 20.
- mochim. Acta 16, 201 (1959). This work was supported by NSF grant GA 10855, by the Petroleum Research Fund (administered by the American Chemical Society), and the Minnesota Geological Sur-vey. H. G. Thode arranged for the first sul-fur isotope determination. The Economic Geology Research Unit of the University of the Witwatersrand, Johannesburg, provided facilities, and its director, D. A. Pretorius, provided valuable help. 23 November 1970