

THE NUMBER OF STANDARD PLANTS PRODUCED FROM
SEED OF OPEN-POLLINATED DWARF PLANTS

Date Ripe Fruit Gathered	Number of Plants Grown	Number of Standard Plants	Number of Dwarf Plants	Per Cent. of Standard Plants
August 10	935	28	907	2.99
" 21	61	0	61	0.00
" 27	128	1	127	0.78
September 4...	51	1	50	1.96
" 22....	995	13	982	1.31
Total	2,170	43	2,127	1.98

Whether or not cross-pollination is caused by wind or insects is not known, although no large insects, such as bees, were seen to visit the plants. Moreover, tomato pollen is dry and seems better adapted to wind transportation. This could be easily tested by screening the dwarf plants. This would not preclude the possibility of cross-pollination by thrips.

Flowers which are bagged in the bud stage and left undisturbed usually do not set fruit. Jarring the plant while the anthers are dehiscing generally suffices to cause pollination. Tomatoes in greenhouses do not set fruit well unless artificially pollinated.

It seems from this evidence that the tomato is naturally only slightly cross-fertilized. Some external agency, however, is generally needed for self-pollination as well as for cross-pollination.

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At the January meeting of the society held January 7, Professor J. A. Miller, of Swarthmore College, read a paper on "The Determination of the Distances of Stars from Us."

He sketched the attempts of Copernicus, Tycho, Braché, Bradley and Sir William Herschel to find a sensible stellar parallax. Perhaps the chief reason for desiring a stellar parallax at that time was that it would establish the truth of the Copernican system upon observational rather than theoretical evidence.

Although these men failed in their attempts to determine a parallax it was while making observa-

tions for that that Bradley discovered the aberration of light and Herschel established the fact that a physical connection exists between the components of certain double stars.

It was 300 years after Copernicus, more than a century after Bradley and a half century after Herschel before the first sensible parallax of a star was actually found when Bessel found the parallax of 61 Cygni and Henderson a parallax of α Centauri. Bessel completed his observations in 1840 and although astronomers have been working assiduously ever since, reliable parallaxes of only about 400 stars have been determined.

At present eight American observatories are working at the problem under the direction of a committee appointed by the American Astronomical Society. Most of these observatories are applying the photographic method devised by Pritchard, of Oxford. This method has since been refined and improved by various men, most notably perhaps by Schlesinger.

The Sproul Observatory of Swarthmore College is one of the eight observatories mentioned above and is spending most of the energies of its staff in that direction. They have determined in all 46 parallaxes. The program contains:

1. All visual binaries whose orbits are well determined.
2. Those visual binaries, the data concerning which leads us to believe their orbits will be determined in the not very distant future.
3. Some spectroscopic binaries.
4. Some stars of large proper motion.
5. Some stars whose hypothetical parallax is large.
6. Other objects.

Classes 1, 2, 3, receive most attention and the measurements and reductions of 13 stars of Class 1 have been completed.

Though no generalization could be made from so small a number of stars as this, yet so far as can be judged from these 13 stars, the orbits of the binaries are comparable in magnitude to the orbits of the planets. The greatest distance between two components of any double star in this list (τ Cygni) being 32 astronomical units, and the least being four astronomical units for 85 Pegasi.

The combined masses of the two components average larger than the sun. The largest mass being of Lalande 9091, which is 48 times the sun and the smallest being 20 Persei which is 0.26 that of the sun.